

# ***AS 4022 Cosmology***

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**Lecture Slides will be on this web page:**

**<http://star-www.st-and.ac.uk/~kdh1/cos/cos.html>**

**Text (intro): Andrew Liddle: Intro to Modern Cosmology**

**(intermediate): Barbara Ryden: Introduction to Cosmology**

**Dan Maoz: Astrophysics in a Nutshell**

**(advanced): John Peacock: Cosmological Physics**

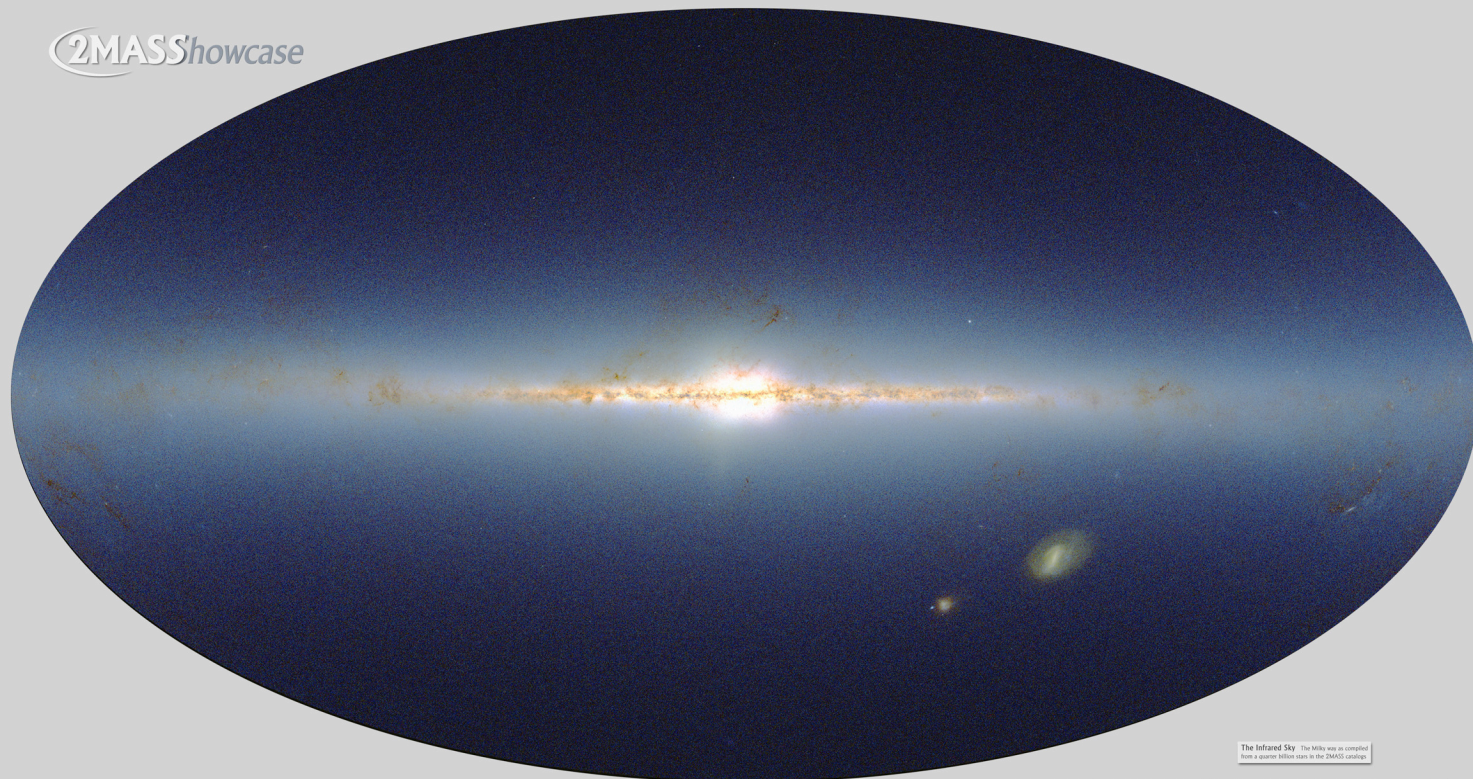
**Web Lecture Notes: John Peacock, Ned Wright**

# ***Lecture 1***

## ***Review / Overview***

# *Why Study Cosmology?*

- **Fascinating questions:**
  - Birth, life, destiny of our Universe
  - Hot Big Bang --> ( 75% H, 25% He ) observed in stars!
  - Formation of structure ( galaxies ... )
- **Technology -> much recent progress:**
  - Precision cosmology: uncertainties of 50% --> 2%
- **Deep mysteries remain:**
  - Dark Matter? Dark Energy? General Relativity wrong?
- **Stretches your mind:**
  - Curved expanding spaces, looking back in time, ...



***Milky Way***

***Andromeda***

The Infrared Sky - The Milky Way as compiled from a quarter billion stars in the 2MASS catalog.

Two Micron All Sky Survey Image Mosaic; Infrared Processing and Analysis Center/Caltech & University of Massachusetts

***~10<sup>10</sup> stars per  
Galaxy:***



© Philip Perkins 1999

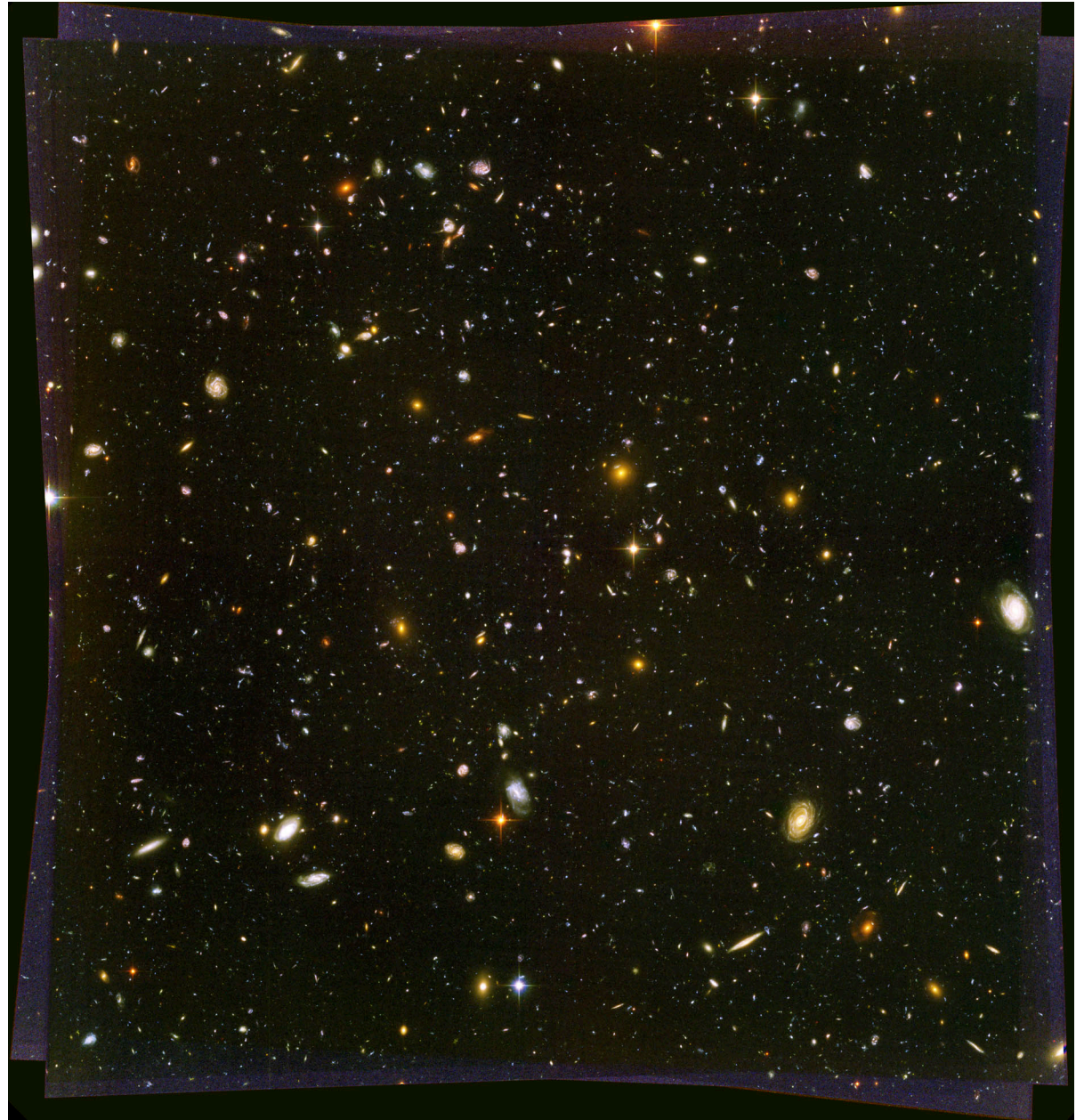
# *Hubble Deep Field:*

At faint magnitudes,  
we see **thousands of  
Galaxies for every  
star !**

$\sim 10^{10}$  galaxies in the  
visible Universe

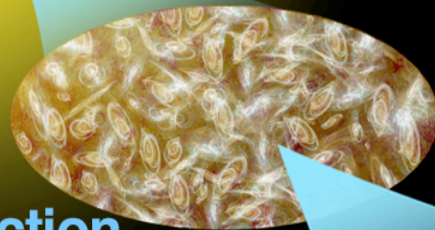
$\sim 10^{10}$  stars per  
galaxy

$\sim 10^{20}$  stars in the  
visible Universe



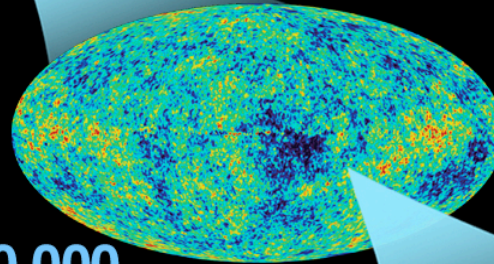
# Cosmology Today

DAWN  
OF  
TIME  
?



tiny fraction  
of a second

inflation



380,000  
years

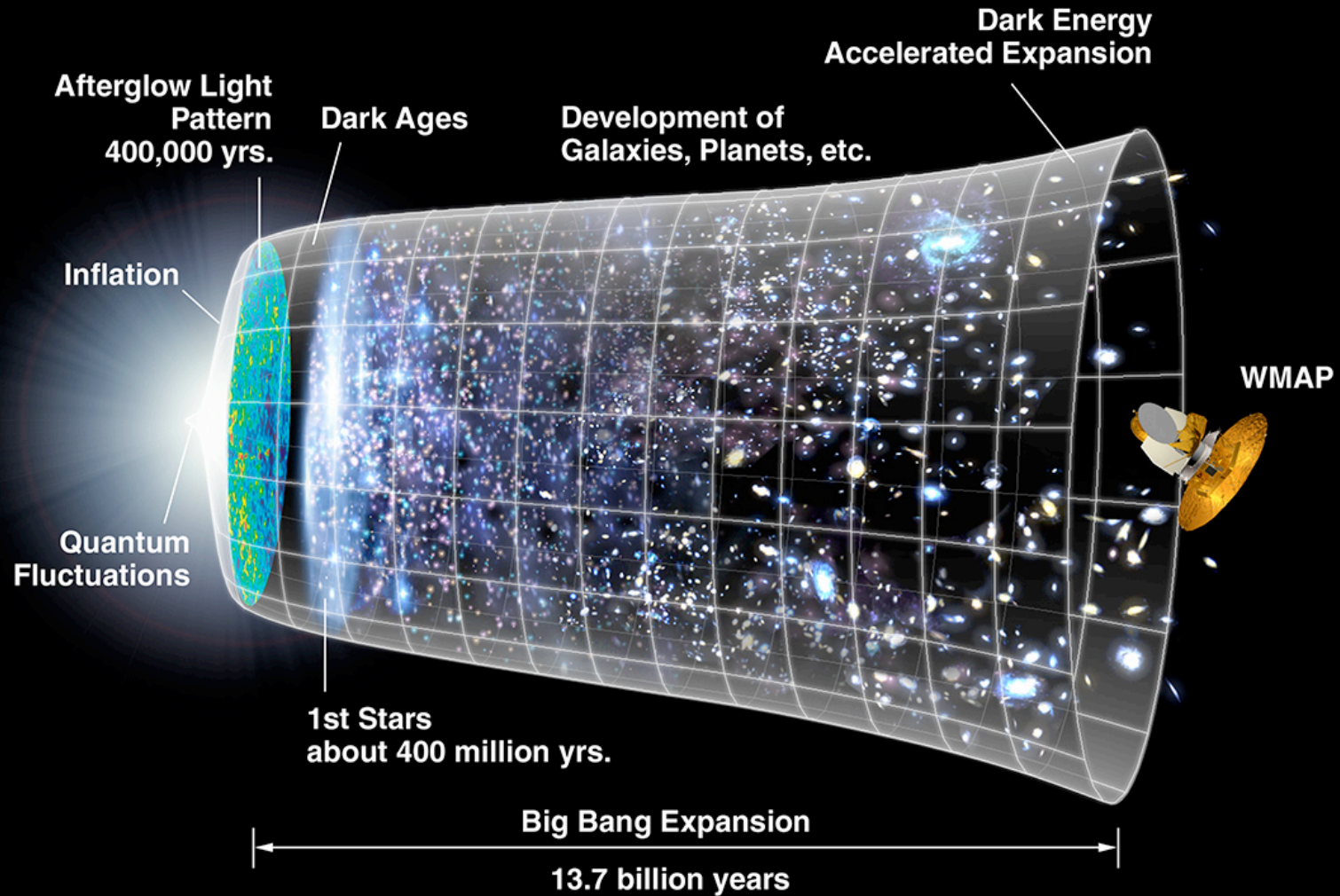
13.7  
billion  
years



# ***1980: Inflation (Alan Guth)***

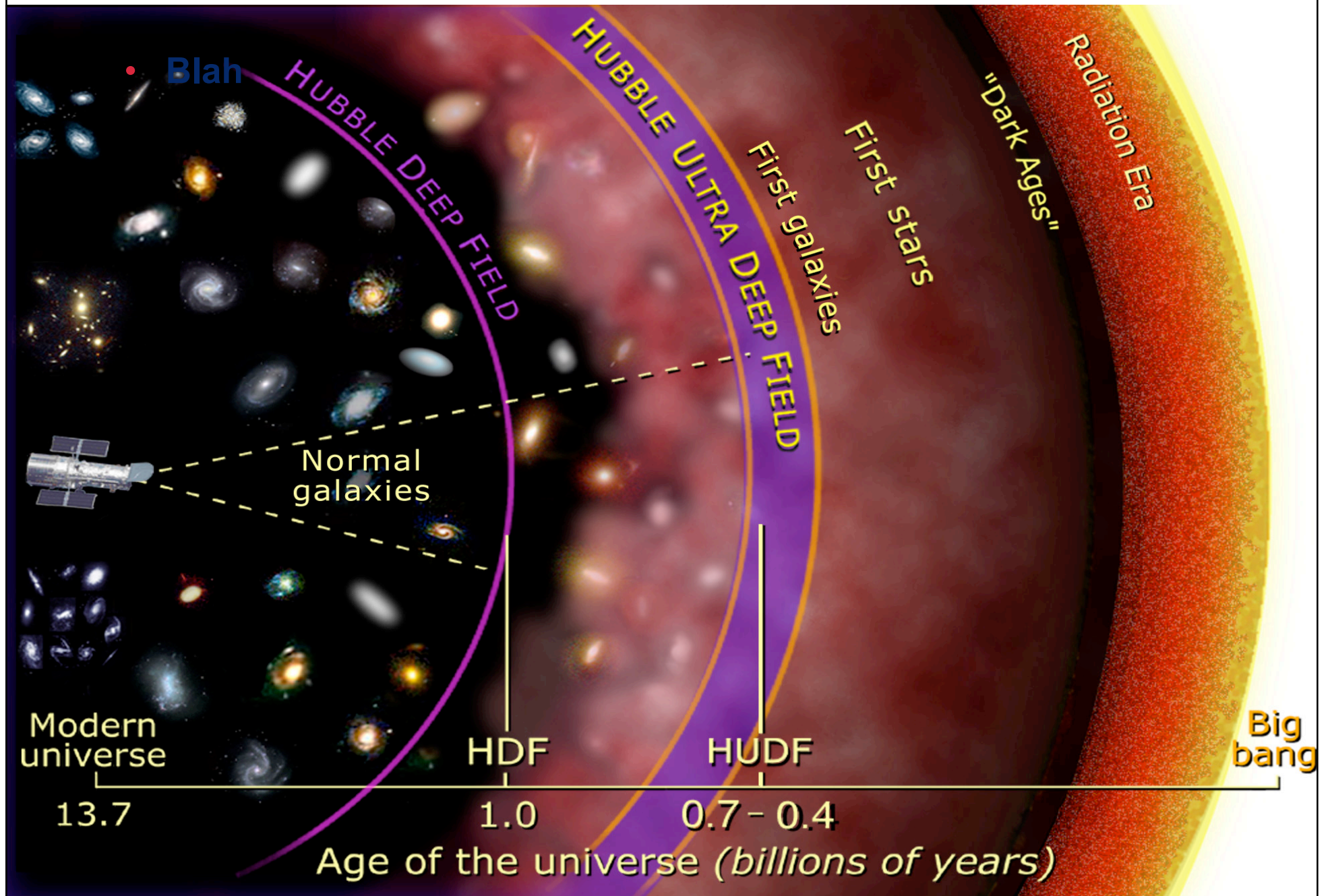
- Universe born from “**nothing**” ?
- A **quantum fluctuation** produces a tiny bubble of “**False Vacuum**”.
- High vacuum energy drives **exponential expansion**, also known as “**inflation.**”
- Universe expands by huge factor in tiny fraction of second, as false vacuum returns to true vacuum.
- Expansion so fast that **virtual particle-antiparticle pairs** get separated to become **real particles and anti-particles.**
- Stretches out all structures, giving a **flat geometry** and uniform  $T$  and  $\rho$ , with **tiny ripples.**
- Inflation launches the **Hot Big Bang!**

# Accelerating/Decelerating Expansion

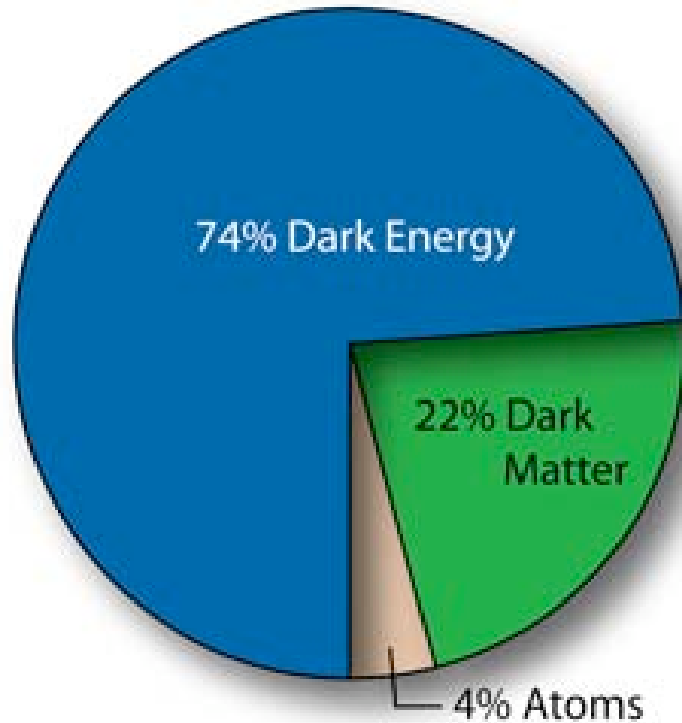




# Looking Back in Time



# ***Current Mysteries***



## ***Dark Matter ?***

*Holds Galaxies together  
Triggers Galaxy formation*

***Dark Energy ?**  
Drives Cosmic Acceleration.*

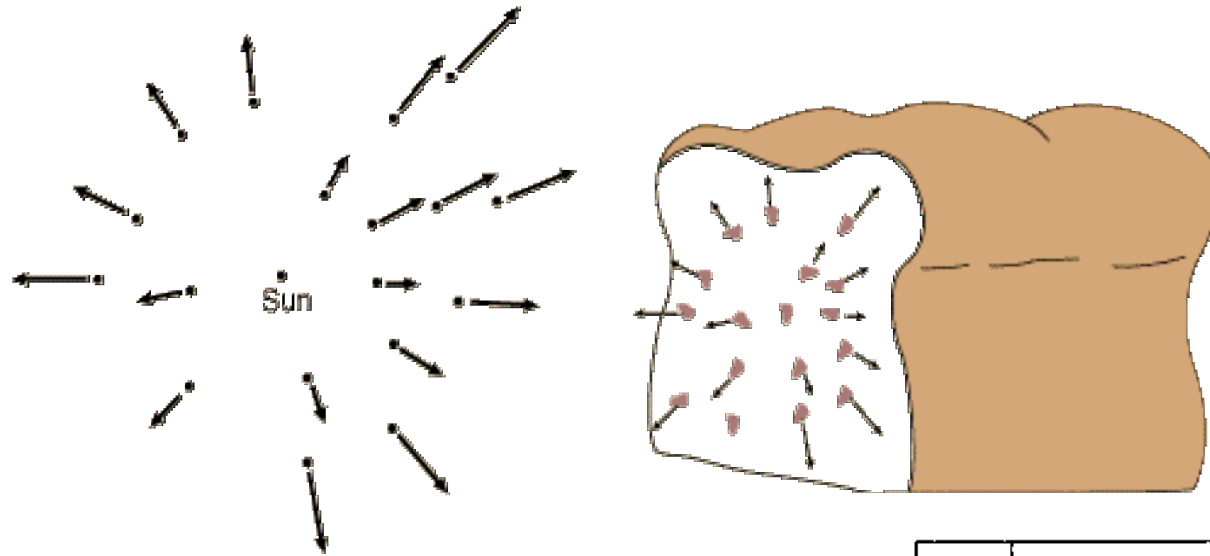
## ***Modified Gravity ?***

*General Relativity wrong ?*

# *Cosmology Milestones*

- 1925 Galaxy redshifts  $\lambda = \lambda_0 (1 + z)$   $V = c z$ 
  - Isotropic expansion. ( Hubble law  $V = H_0 d$  )
  - Finite age. (  $t_0 = 13 \times 10^9$  yr )
- 1965 Cosmic Microwave Background (CMB)
  - Isotropic blackbody.  $T_0 = 2.7$  K
  - Hot Big Bang  $T = T_0 (1 + z)$
- 1925 General Relativity Cosmology Models :
  - Radiation era:  $R \sim t^{1/2}$   $T \sim t^{-1/2}$
  - Matter era:  $R \sim t^{2/3}$   $T \sim t^{-2/3}$
- 1975 Big Bang Nucleosynthesis (BBN)
  - light elements (  $^1\text{H} \dots ^7\text{Li}$  )  $t \sim 3$  min  $T \sim 10^9$  K
  - primordial abundances (75% H, 25% He) as observed!

# *Isotropic Expansion*



Hubble law :

$$V = H_0 d$$

Hubble "constant":

$$H_0 \approx 500 \text{ km s}^{-1} \text{ Mpc}^{-1}$$

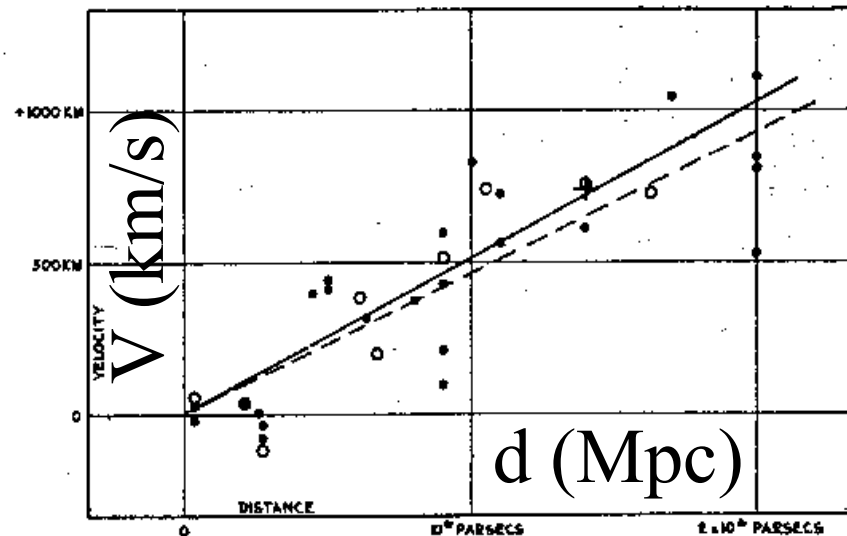


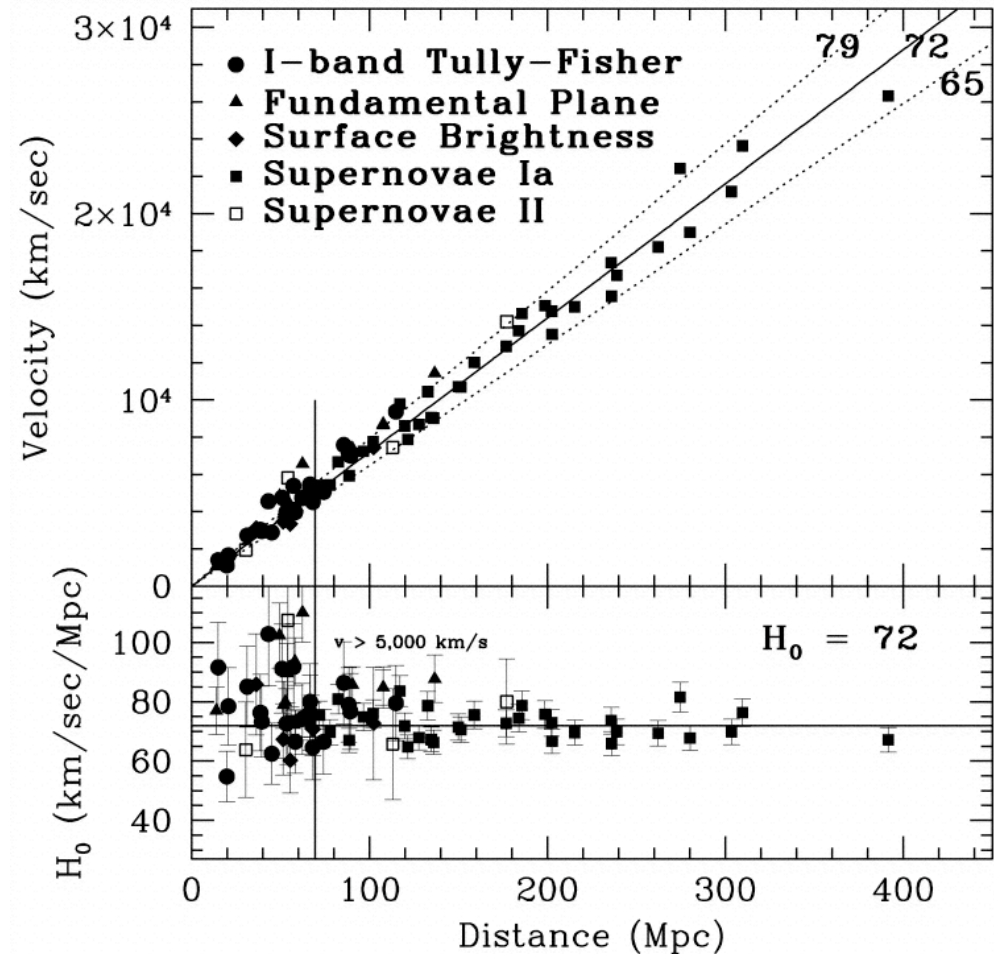
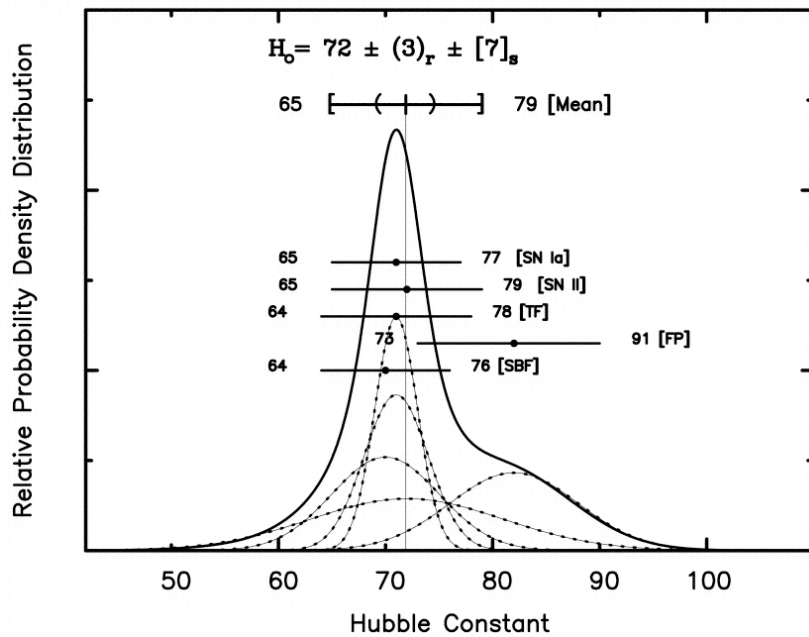
FIGURE 1

WRONG ! Extinction by interstellar dust was not then known, giving incorrect distances.

# $H_0$ from the HST Key Project

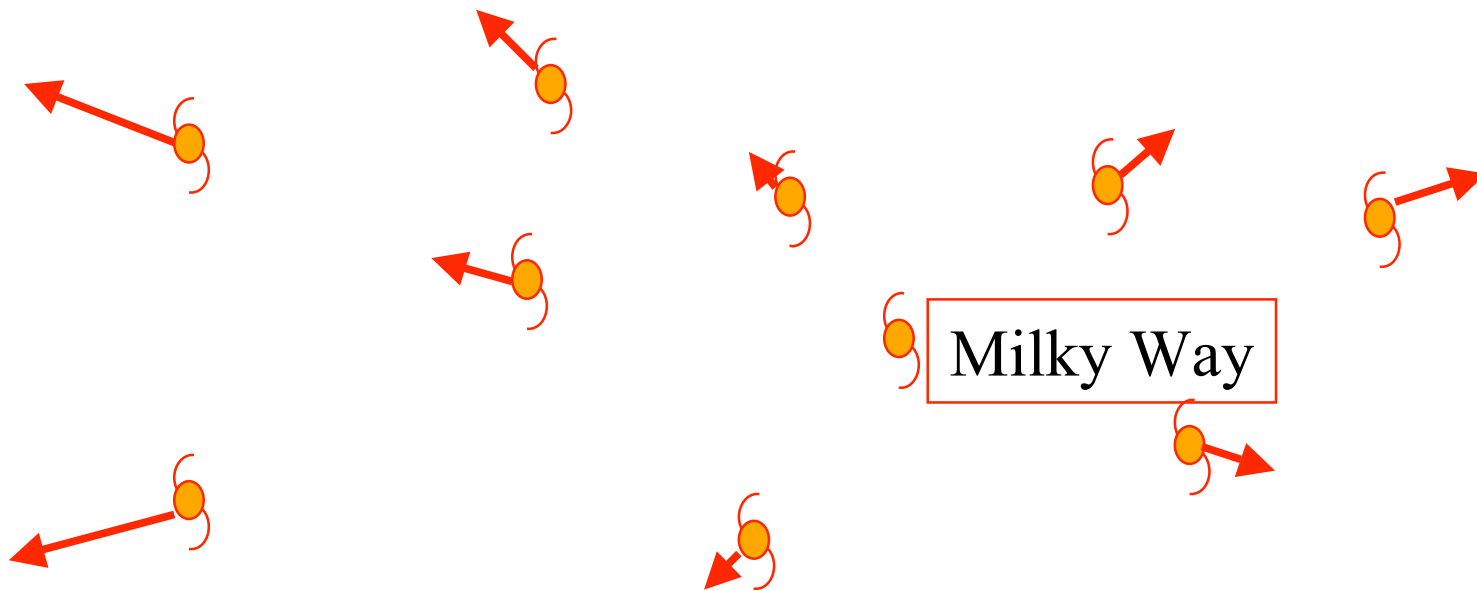
$$H_0 \approx 72 \pm 3 \pm 7 \text{ km s}^{-1} \text{ Mpc}^{-1}$$

Freedman, et al.  
2001 ApJ 553, 47.



# *Universal Expansion*

Hubble's law appears to violate  
The Copernican Principle.  
Are we at a special location?

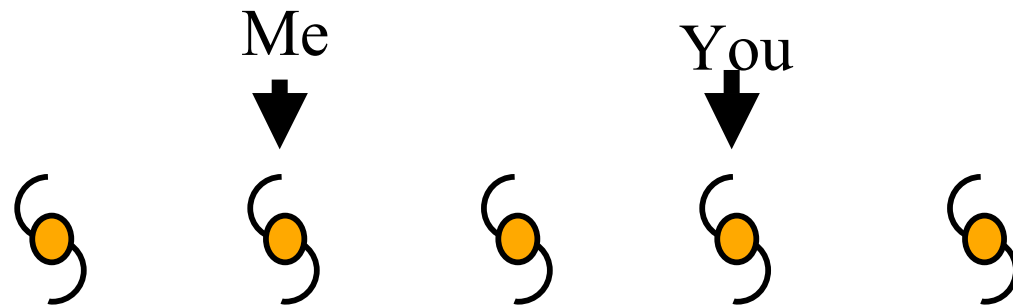


Is everything moving away from us?

# Universal Expansion

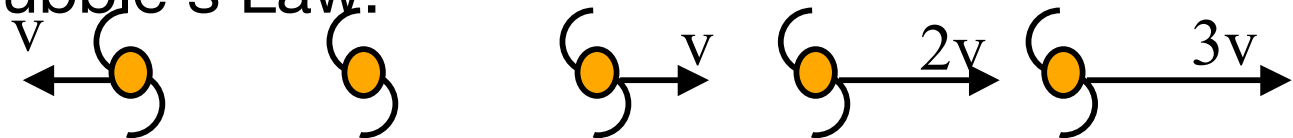
Q : What is so special about our location ?

A : Nothing !

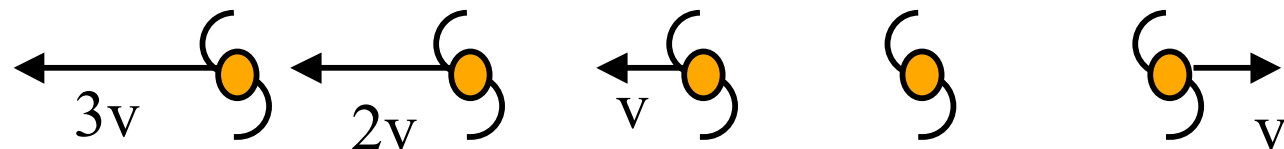


According to Hubble's Law:

I see:



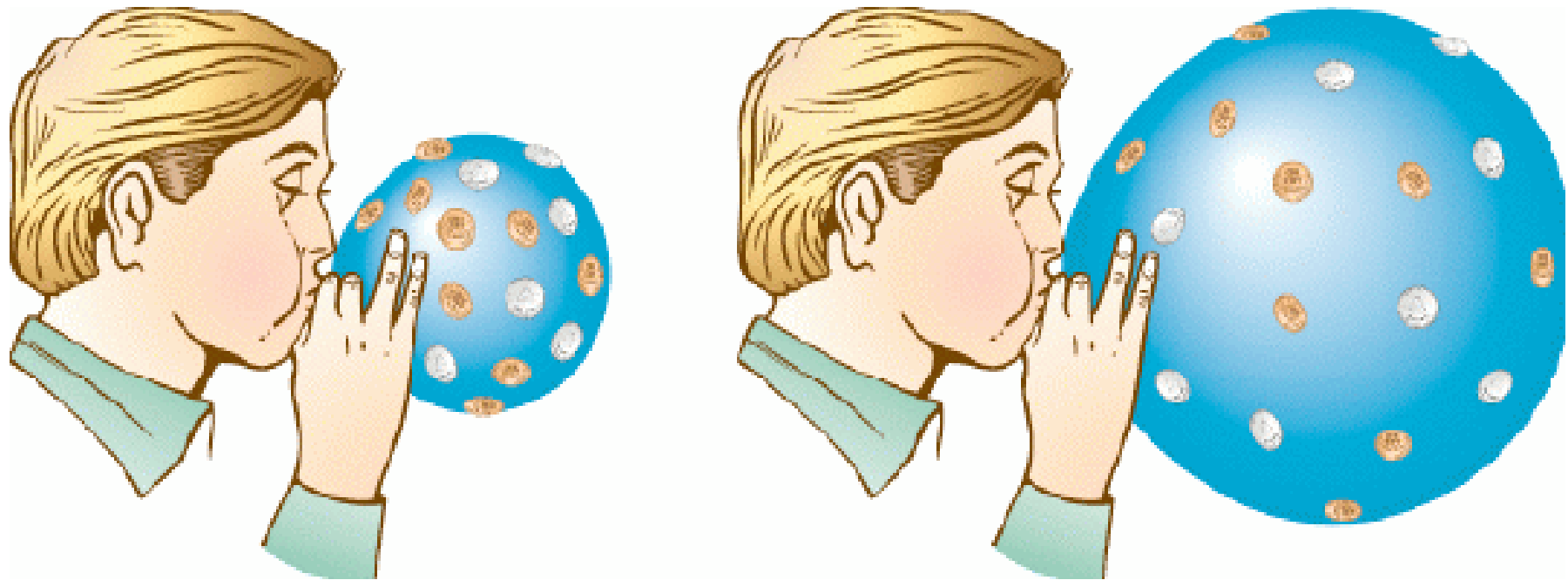
You see:



We all see the same Hubble law expansion.

# *The Universal Expansion*

- An observer in any galaxy sees all other galaxies moving away, with the same Hubble law.
- Expansion (or contraction) produces a centre-less but dynamic Universe.

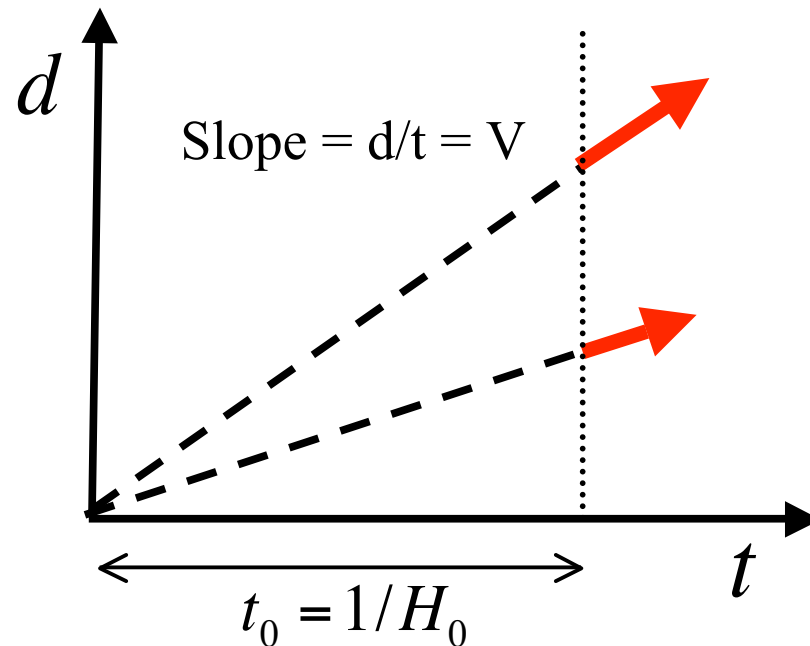




# ***Hubble Law --> Finite age.***

$$V = H_0 d$$

$$t_0 \approx \frac{d}{V} = \frac{1}{H_0} = \left( \frac{1 \text{ Mpc}}{72 \text{ km/s}} \right) \left( \frac{3 \times 10^{19} \text{ km}}{\text{Mpc}} \right) \left( \frac{1 \text{ yr}}{3 \times 10^7 \text{ s}} \right)$$
$$\approx 13 \times 10^9 \text{ yr} = 13 \text{ Gyr.}$$



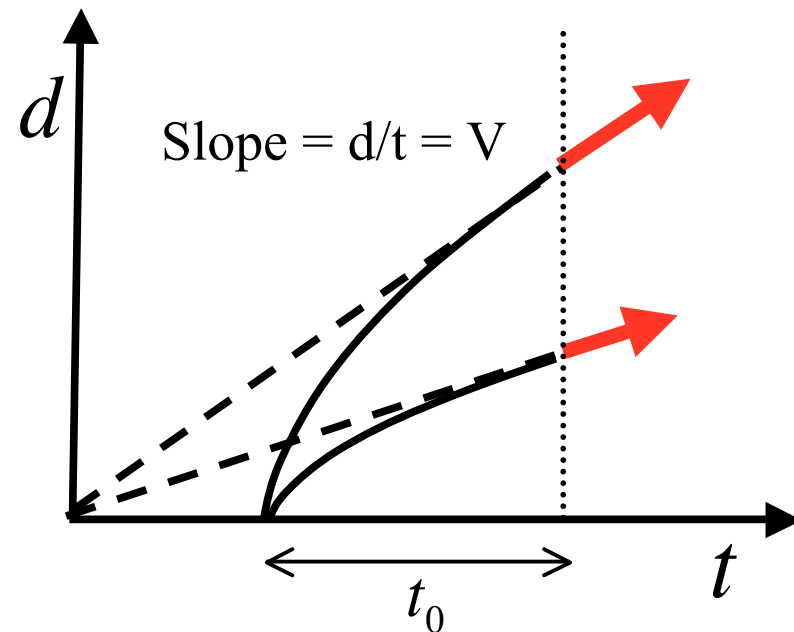
# *Hubble Law --> Finite age.*

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Gravity decelerates:

$$t_0 \approx \frac{2}{3} \frac{1}{H_0} .$$



# *Hubble Law --> Finite age.*

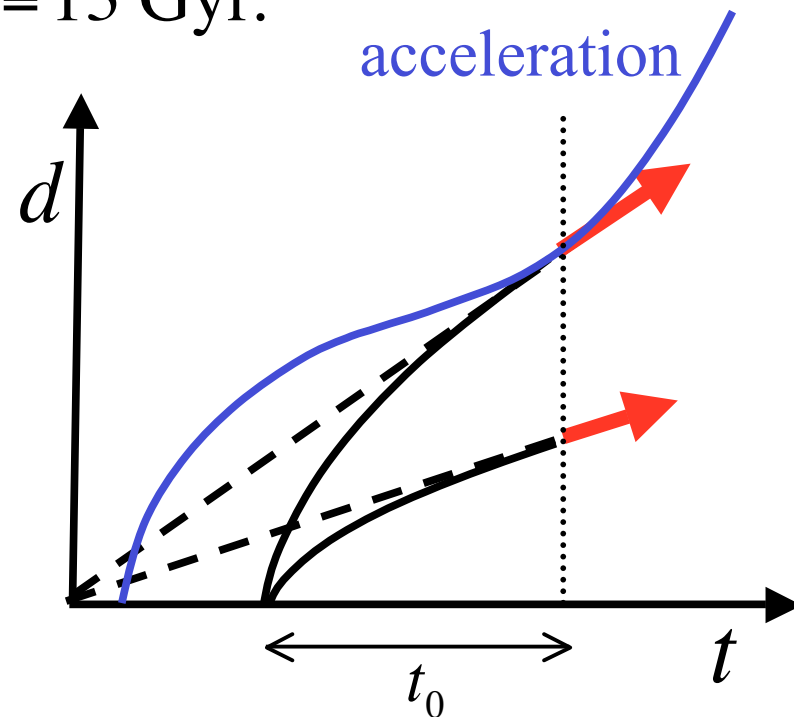
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Gravity decelerates :

Dark Energy accelerates

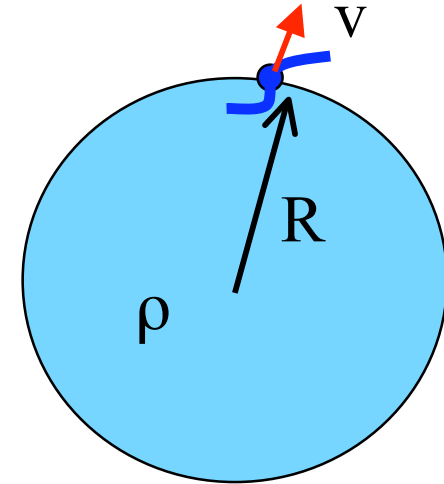
$$t_0 > \frac{2}{3} \frac{1}{H_0} .$$



# Critical Density

- Newtonian analogy:  
escape velocity:

$$V_{esc}^2 = \frac{2GM}{R} = \frac{2G}{R} \left( \frac{4\pi R^3 \rho}{3} \right) = \frac{8\pi G R^2 \rho}{3}$$

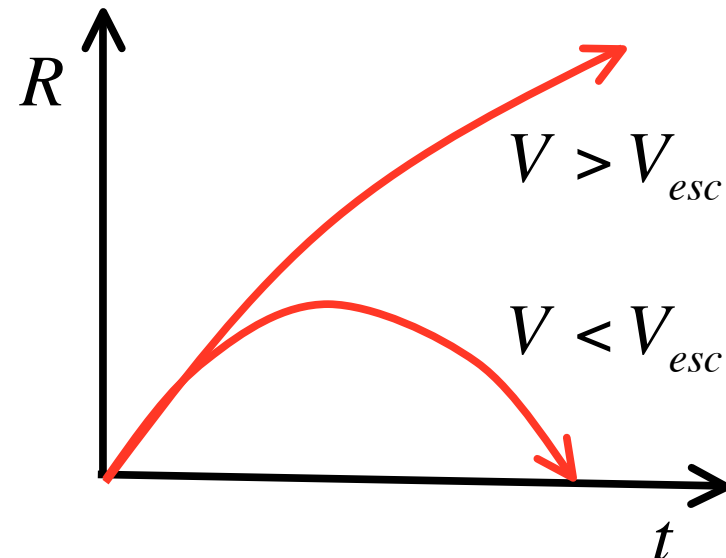


Hubble expansion:

$$V = H_0 R$$

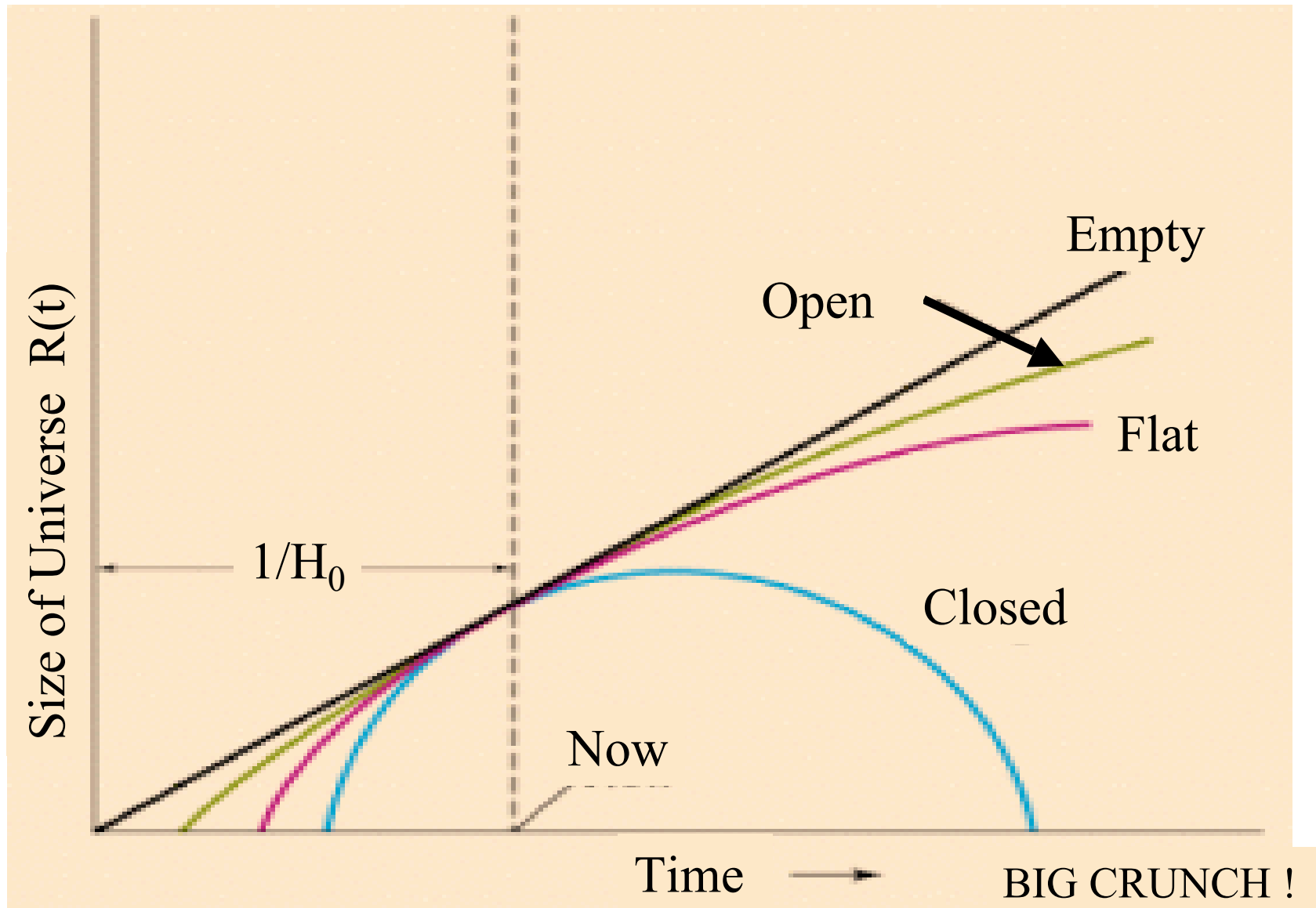
critical density:

$$\left( \frac{V_{esc}}{V} \right)^2 = \frac{8\pi G \rho}{3 H_0^2} = \frac{\rho}{\rho_c}$$



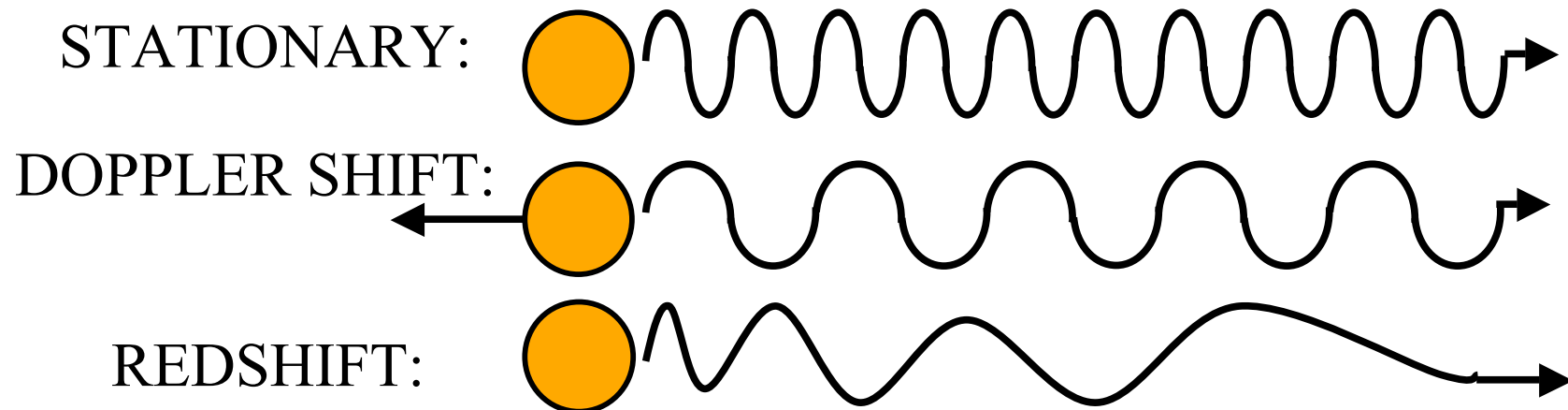
$$\rho_c \equiv \frac{3 H_0^2}{8\pi G} \approx 10^{-26} \text{ kg m}^{-3} \approx \frac{1.4 \times 10^{11} \text{ Msun}}{(\text{Mpc})^3}$$

# Re-collapse or Eternal Expansion ?



# Redshift

- Expansion is a stretching of space.
- The more space there is between you and a galaxy, the faster it appears to be moving away.
- Expansion **stretches the wavelength of light**, causing a galaxy's spectrum to be **REDSHIFTED**:



**REDSHIFT IS NOT THE SAME AS DOPPLER SHIFT**

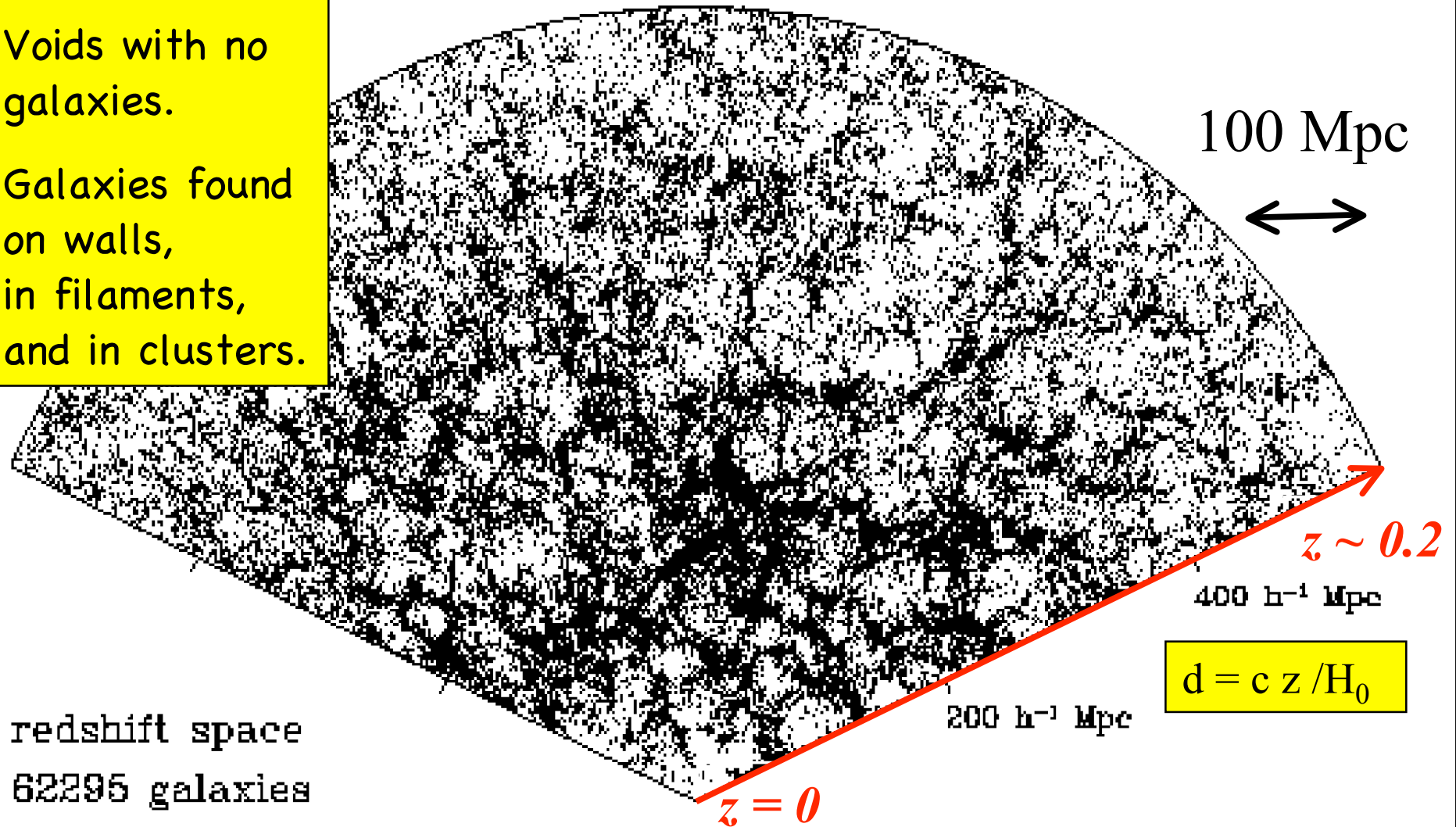
# Galaxy Redshift Surveys

Bubble-like structure:

Voids with no galaxies.

Galaxies found on walls, in filaments, and in clusters.

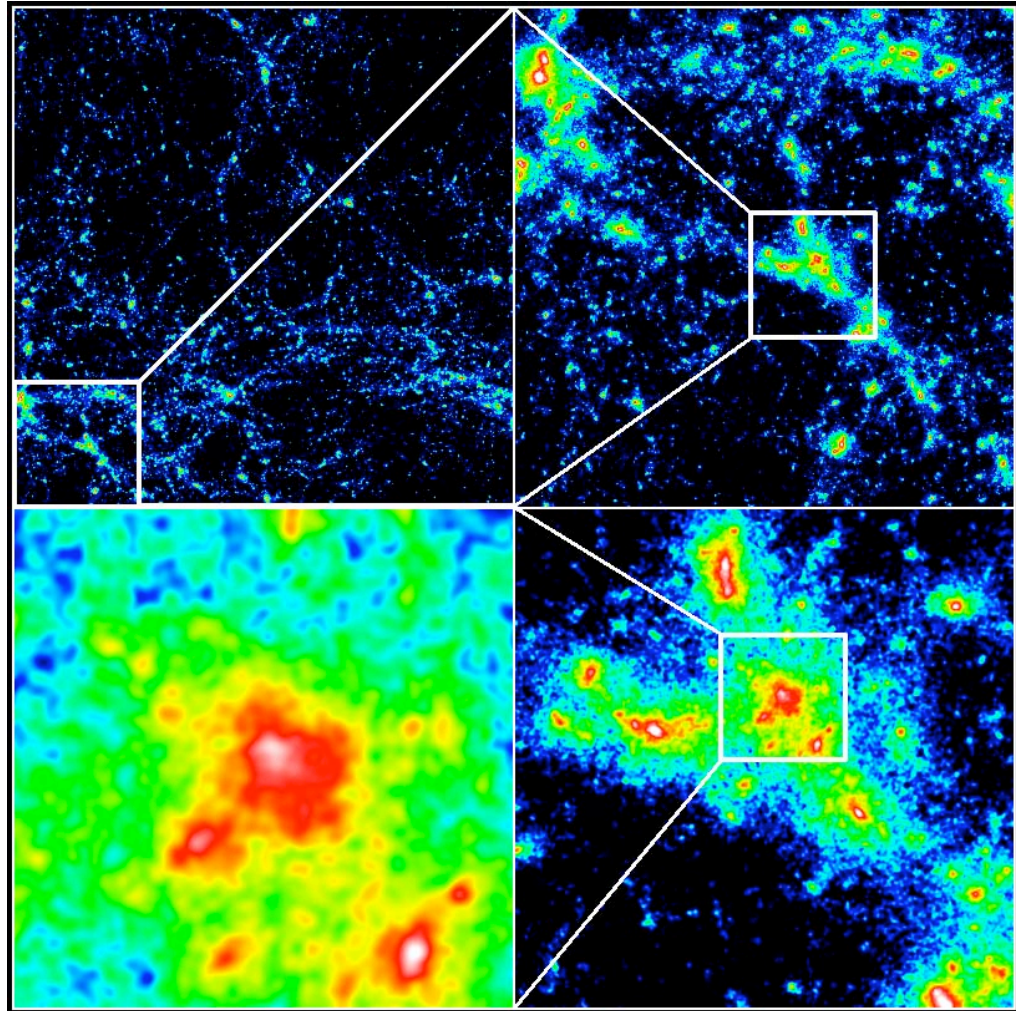
$r' < 17.55$ ,  $d > 2''$ ,  $6^\circ$  slice



redshift space  
62295 galaxies

# The Visible Cosmos: a hierarchy of structure and motion

- Computer simulations of structure formation:





# *Cosmological Models*

Einstein's gravity theory (General Relativity)

Assume Universe filled with uniform density fluid.

[ OK on large scales  $> 100$  Mpc ]

Density:  $\rho = \Omega \rho_c$       Energy density:  $\varepsilon = \rho c^2$

Critical density:  $\rho_c \equiv \frac{3 H_0^2}{8\pi G} \approx 10^{-26} \text{kg m}^{-3} \approx \frac{1.4 \times 10^{11} \text{Msun}}{(\text{Mpc})^3}$

3 components:

1. *Radiation*       $\Omega_R \approx 5 \times 10^{-5}$

2. *Matter*       $\Omega_M \sim 0.3 \left\{ \begin{array}{ll} \text{“Dark Matter”} & \text{baryons} \\ \Omega_D \sim 0.26 & \Omega_B \sim 0.04 \end{array} \right.$

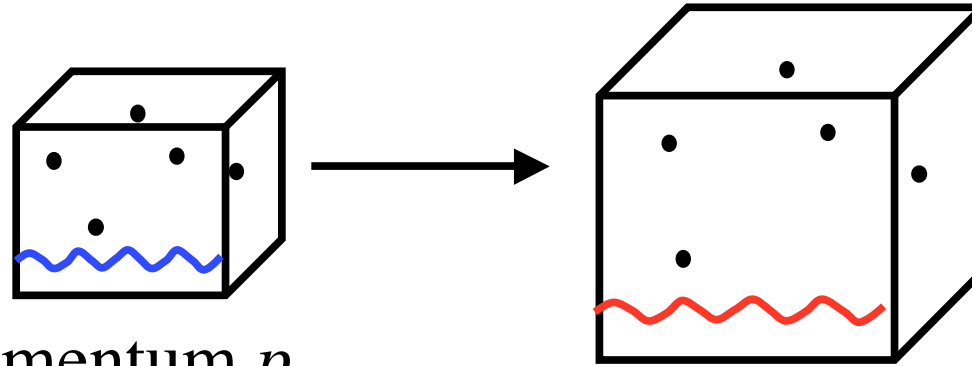
3. *“Dark Energy”*       $\Omega_\Lambda \sim 0.7$

Total       $\Omega = \Omega_R + \Omega_M + \Omega_\Lambda = 1$

*Only ~4% is matter  
as we know it!*

# *Energy Density of expanding box*

volume  $R^3$   
 $N$  particles



particle mass  $m$       momentum  $p$

energy  $E = h\nu = \sqrt{m^2 c^4 + p^2 c^2} = m c^2 + \frac{p^2}{2m} + \dots$

**Cold Matter:** ( $m > 0, p \ll mc$ )

$$E \approx m c^2 = \text{const}$$

$$\epsilon_M \approx \frac{N m c^2}{R^3} \propto R^{-3}$$

**Radiation:** ( $m = 0$ )

**Hot Matter:** ( $m > 0, p \gg mc$ )

$\lambda \propto R$  (wavelengths stretch):

$$E = h\nu = \frac{hc}{\lambda} \propto R^{-1}$$

$$\epsilon_R = \frac{N h\nu}{R^3} \propto R^{-4}$$

### 3 Eras: radiation...matter...vacuum

radiation:  $\rho_R \propto R^{-4}$

matter:  $\rho_M \propto R^{-3}$

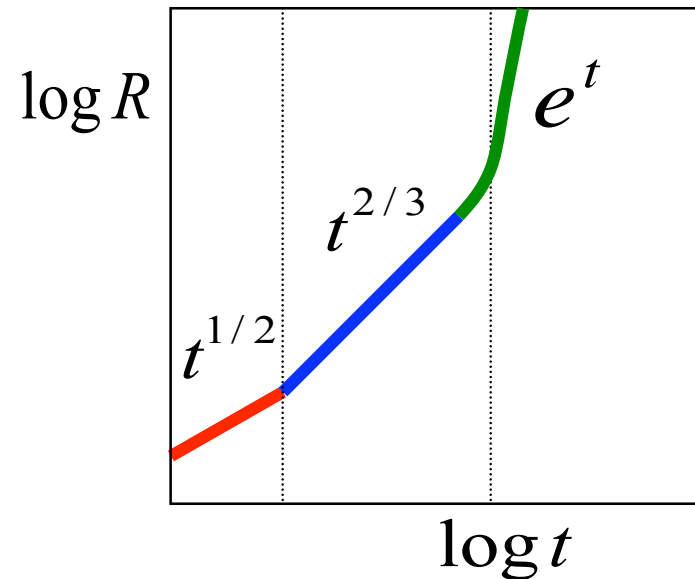
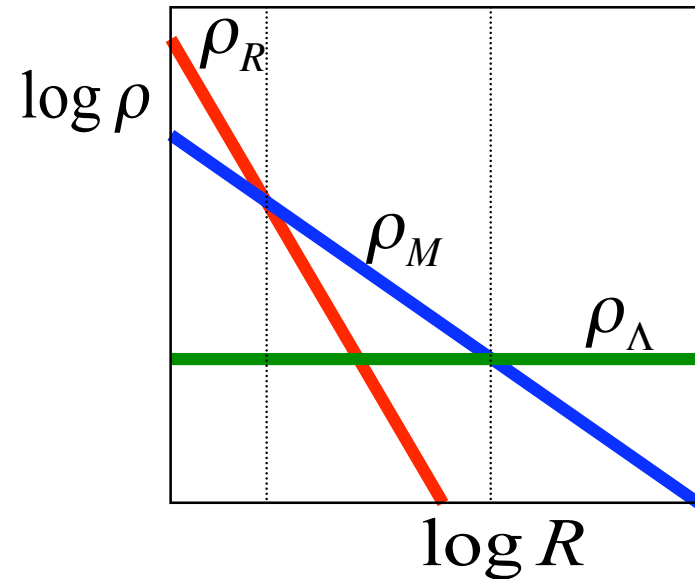
vacuum:  $\rho_\Lambda = \text{const}$

$$a \equiv \frac{R}{R_0} = \frac{1}{1+z}$$

$$\rho = \frac{\rho_{R,0}}{a^4} + \frac{\rho_{M,0}}{a^3} + \rho_\Lambda$$

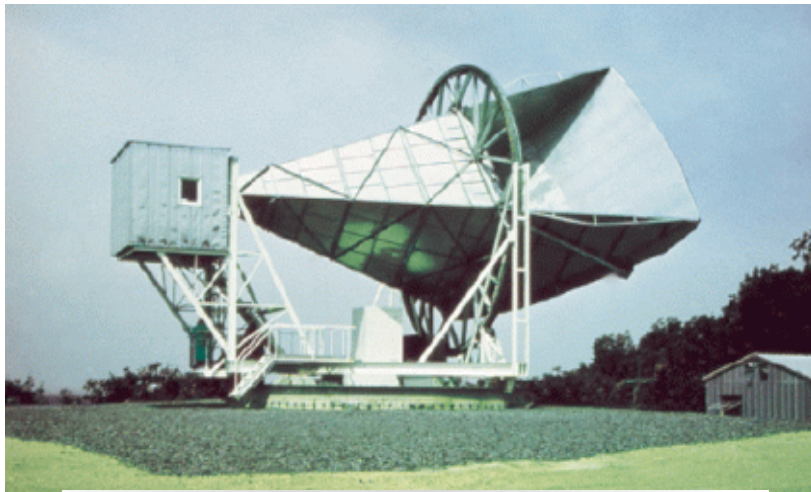
$$\rho_R = \rho_M \text{ at } a \sim 10^{-4} \quad t \sim 10^4 \text{ yr}$$

$$\rho_M = \rho_\Lambda \text{ at } a \sim 0.7 \quad t \sim 10^{10} \text{ yr}$$

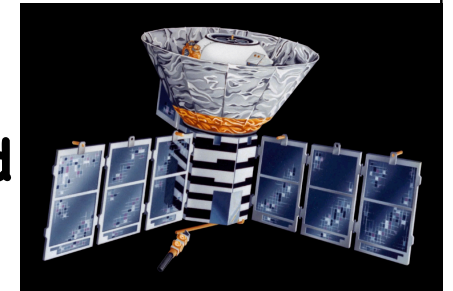


# ***Cosmic Microwave Background***

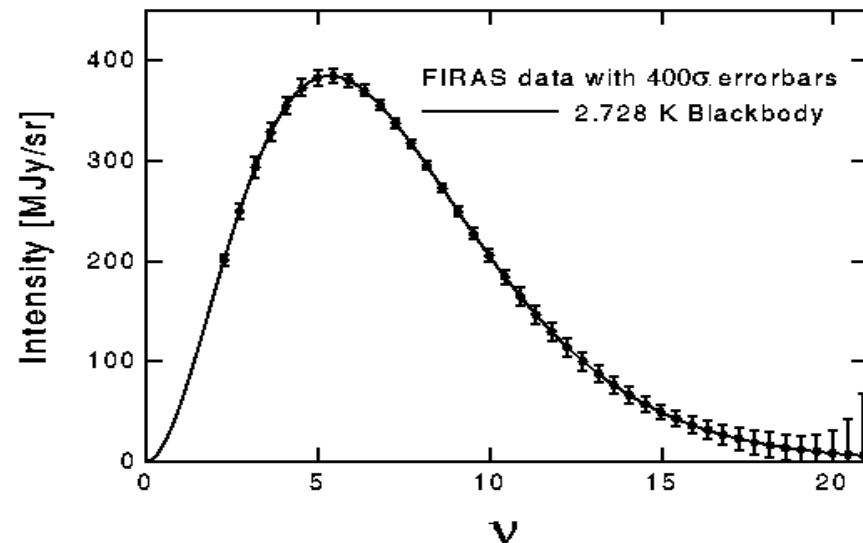
- CMB predicted by Gamov in 1948.  
Discovered by Penzias and Wilson in 1965.



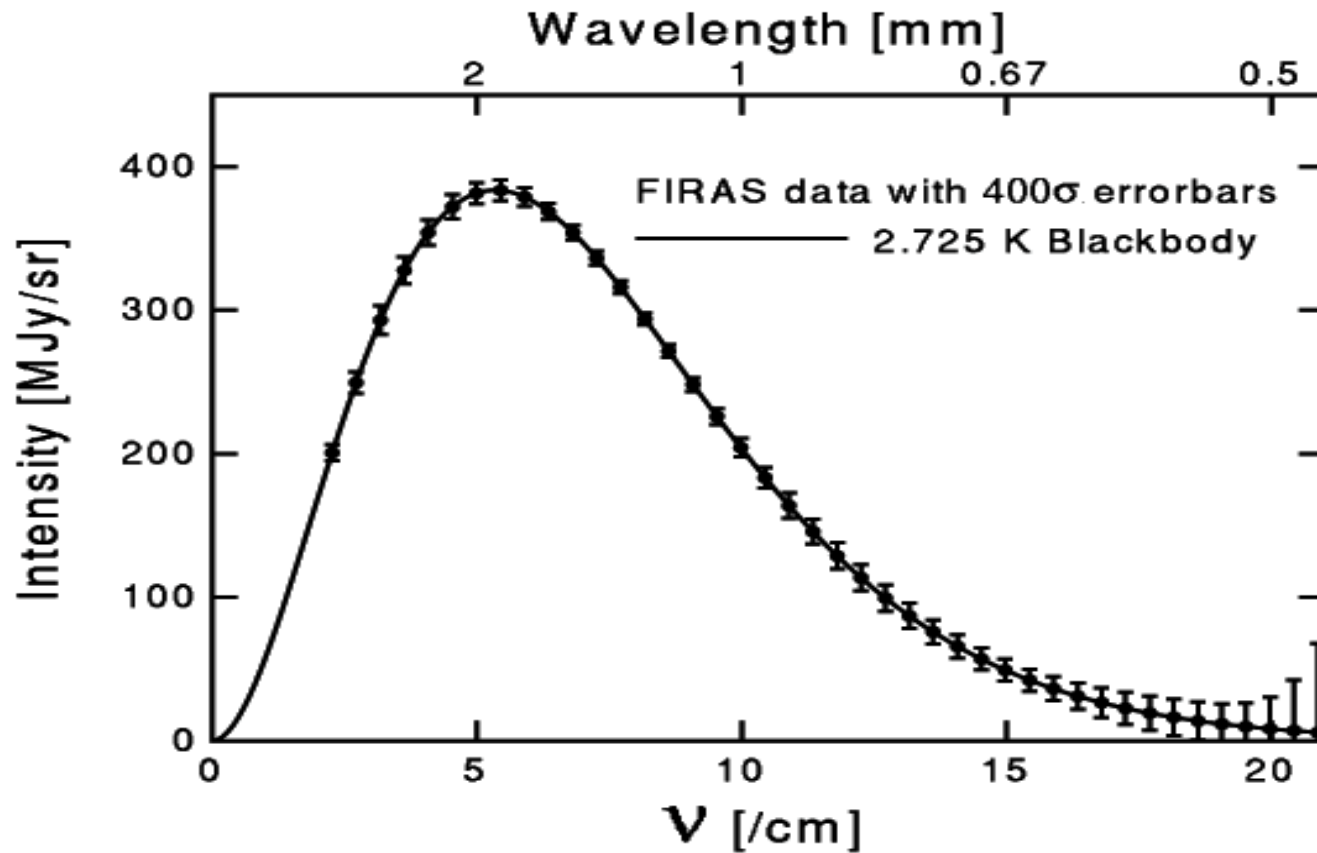
1992 NASA -  
**COBE**  
**CO**smic **B**ackground  
**E**xplorer



**A perfect Blackbody!**



# COBE spectrum of CMB



**A perfect Blackbody !**

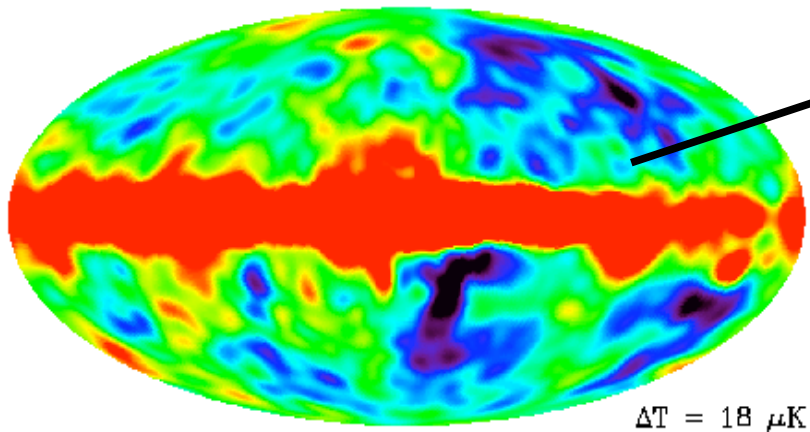
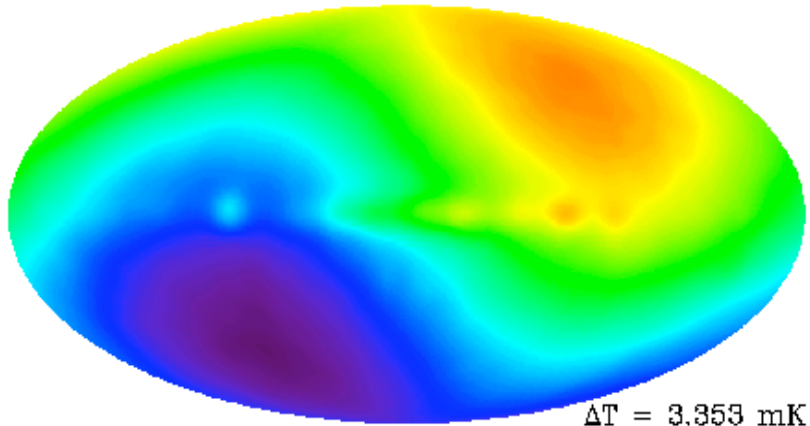
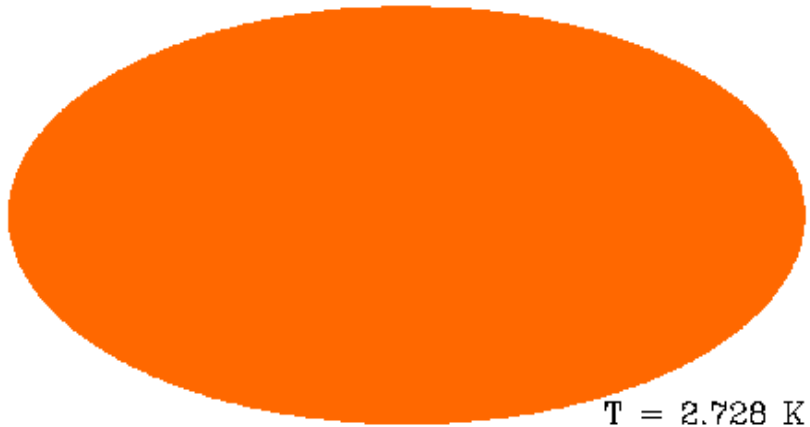
**No spectral lines -- strong test of Big Bang.  
Expansion preserves the blackbody spectrum.**

$$T(z) = T_0 (1+z) \quad T_0 \sim 3000 \text{ K} \quad z \sim 1100$$

# Cosmic Microwave Background

Almost isotropic

$$T = 2.728 \text{ K}$$



Dipole anisotropy

$$\frac{V}{c} = \frac{\Delta\lambda}{\lambda} = \frac{\Delta T}{T} \approx 10^{-3}$$

Our velocity:

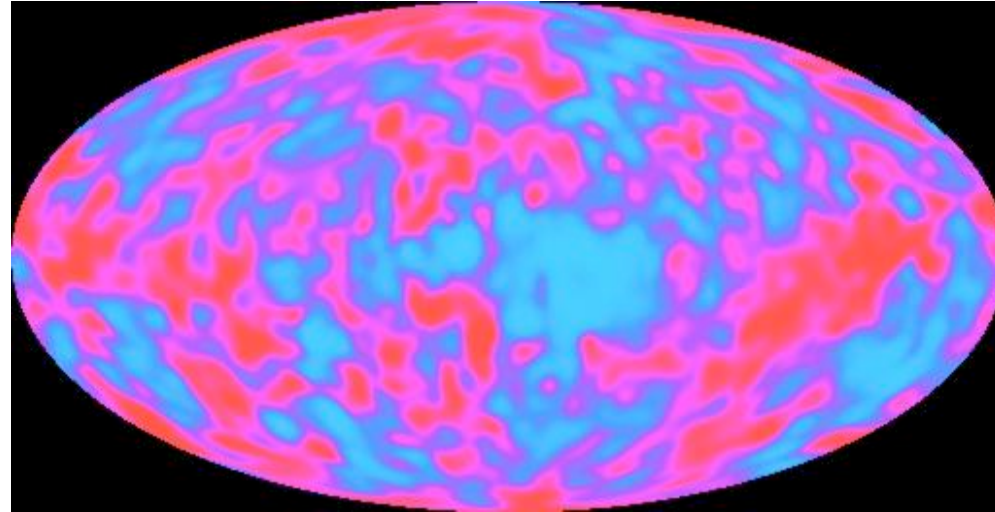
$$V \approx 600 \text{ km/s}$$

Milky Way sources

$$+ \text{ anisotropies } \frac{\Delta T}{T} \sim 10^{-5}$$

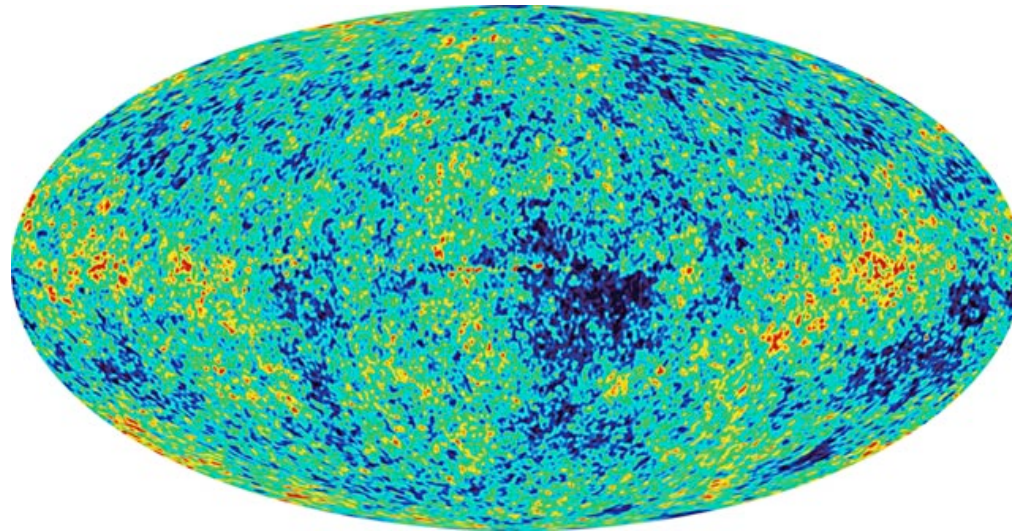
# ***CMB Anisotropies***

COBE  
1994



$$\frac{\Delta T}{T} \sim 10^{-5}$$

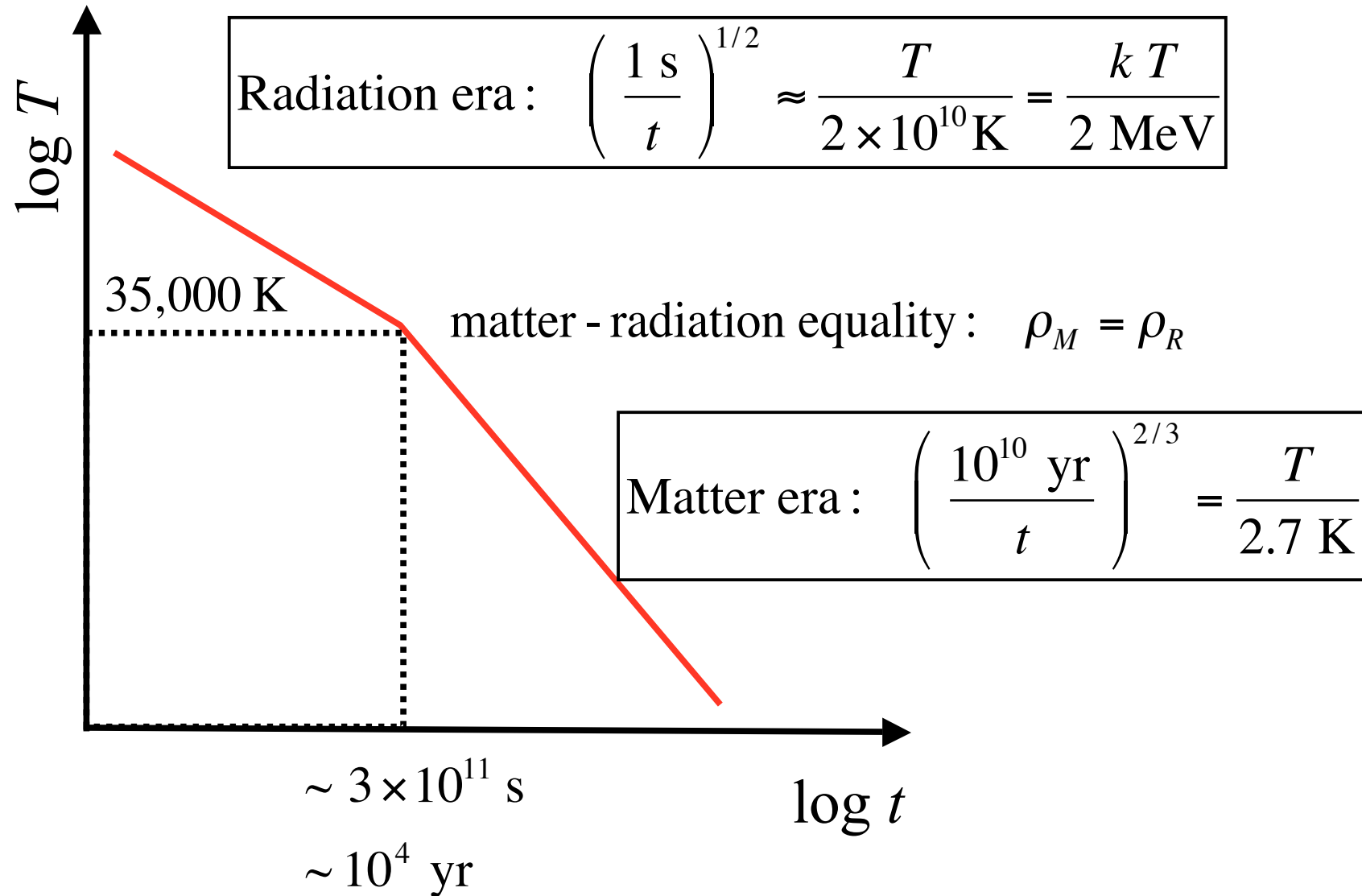
WMAP  
2004



$$\Delta\theta \sim 1^\circ$$

Snapshot of Universe at  $z = 1100$   
Seeds that later form galaxies.

# Cooling History: $T(t)$





# 1975: Big Bang Nuclear Fusion

Big Bang + 3 minutes

$T \sim 10^9 \text{ K}$

First atomic nuclei forged.

Calculations predict:

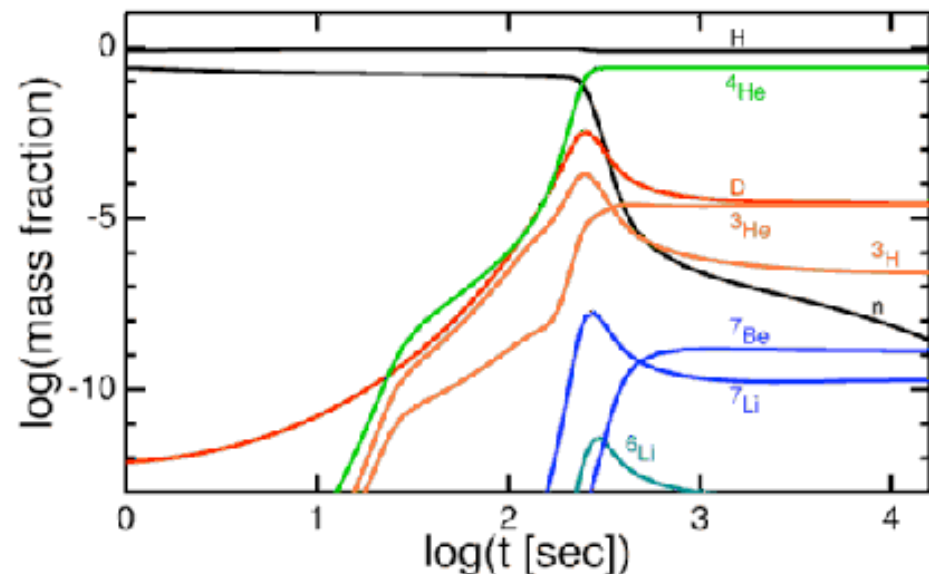
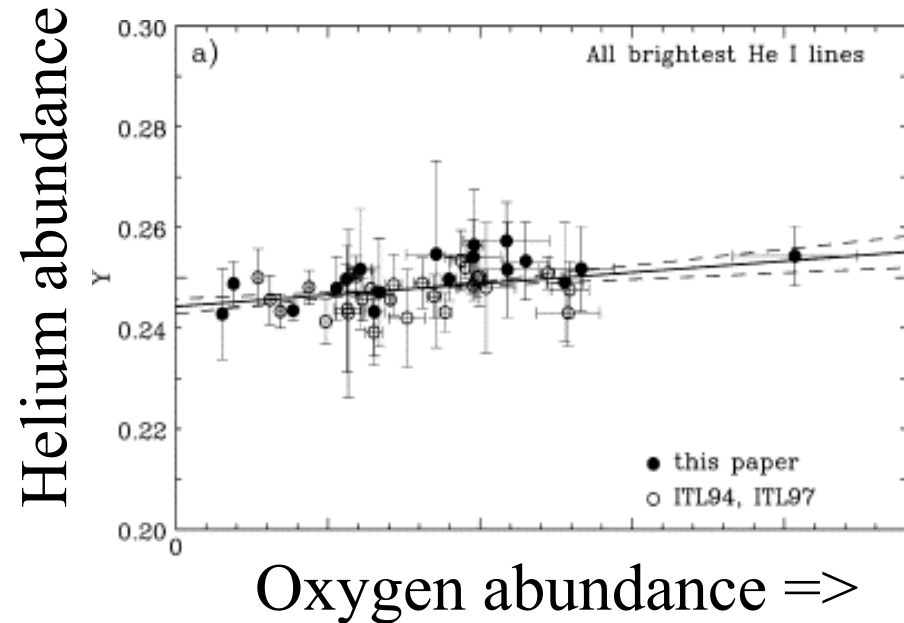
75% H and 25% He

**AS OBSERVED !**

+ traces of light elements

D,  $^3\text{H}$ ,  $^3\text{He}$ ,  $^7\text{Be}$ ,  $^7\text{Li}$

$\Rightarrow$  normal matter only 4% of critical density.



# 1998: *Supernova Cosmology*

- Do galaxies at VERY large distances have the same distance/velocity relationship as the Hubble Law?
- Has the rate of expansion changed?
- SN Ia as “standard candles”: same maximum  $L$
- Search lots of galaxies for SN Ia: very bright

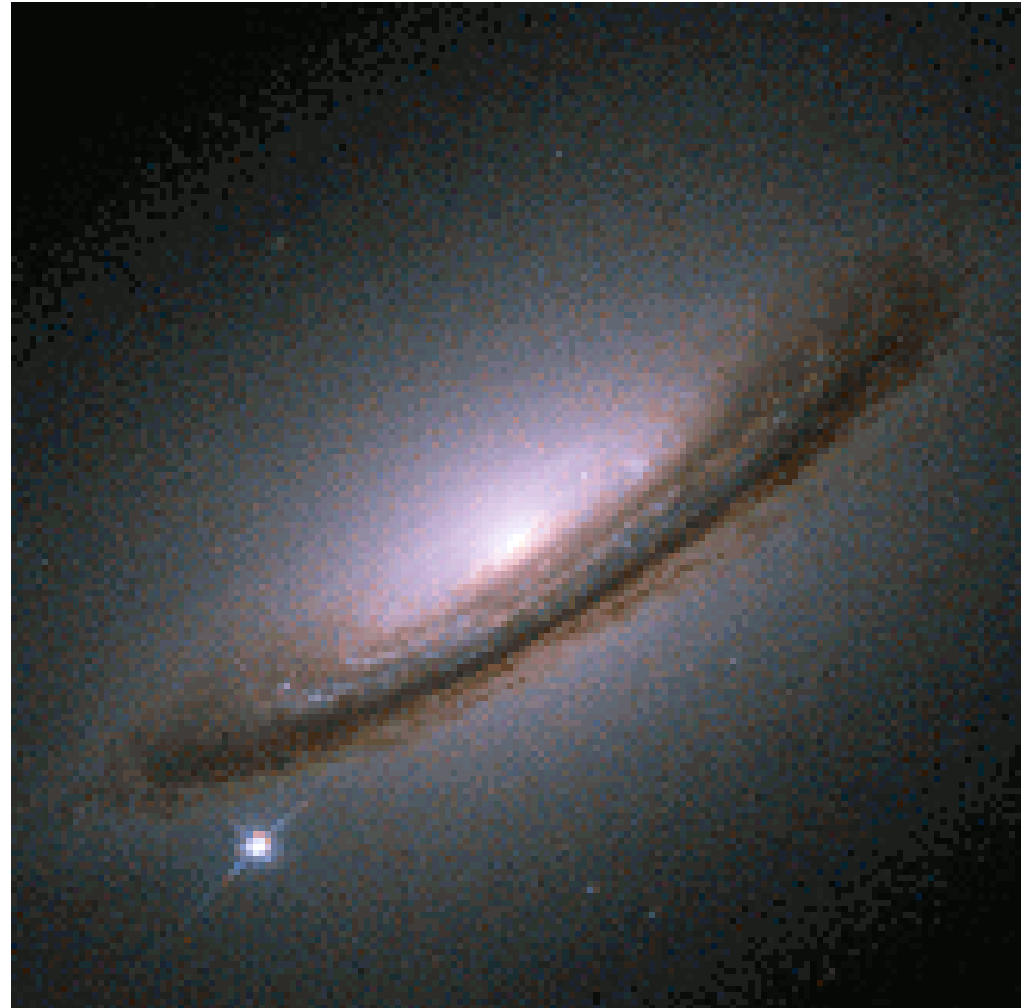
# *SN Type Ia in Virgo Galaxy NGC 4526*

**Supernova outshines the entire galaxy, but only for a month or so.**

Type II -- massive stars ( $M > 8 M_{\text{SUN}}$ ) explode at end of life.

Type Ia -- white dwarf in a binary system accretes mass, collapses when  $M_{\text{WD}} = 1.4 M_{\text{SUN}}$ .

Good “standard bombs”.



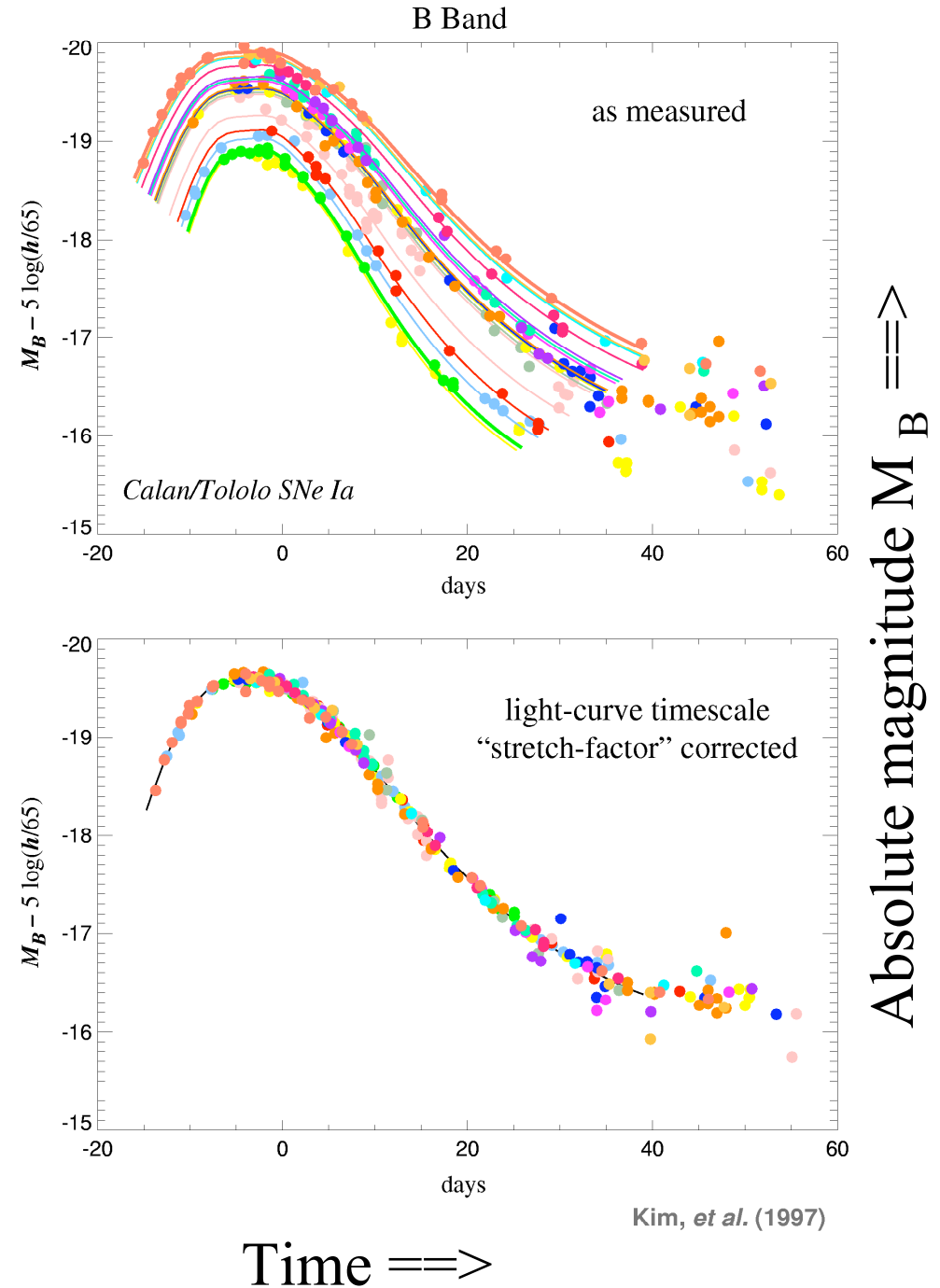
Calibrate SN distances using HST to see Cepheids in Virgo galaxies.

# Calibrating “Standard Bombs”

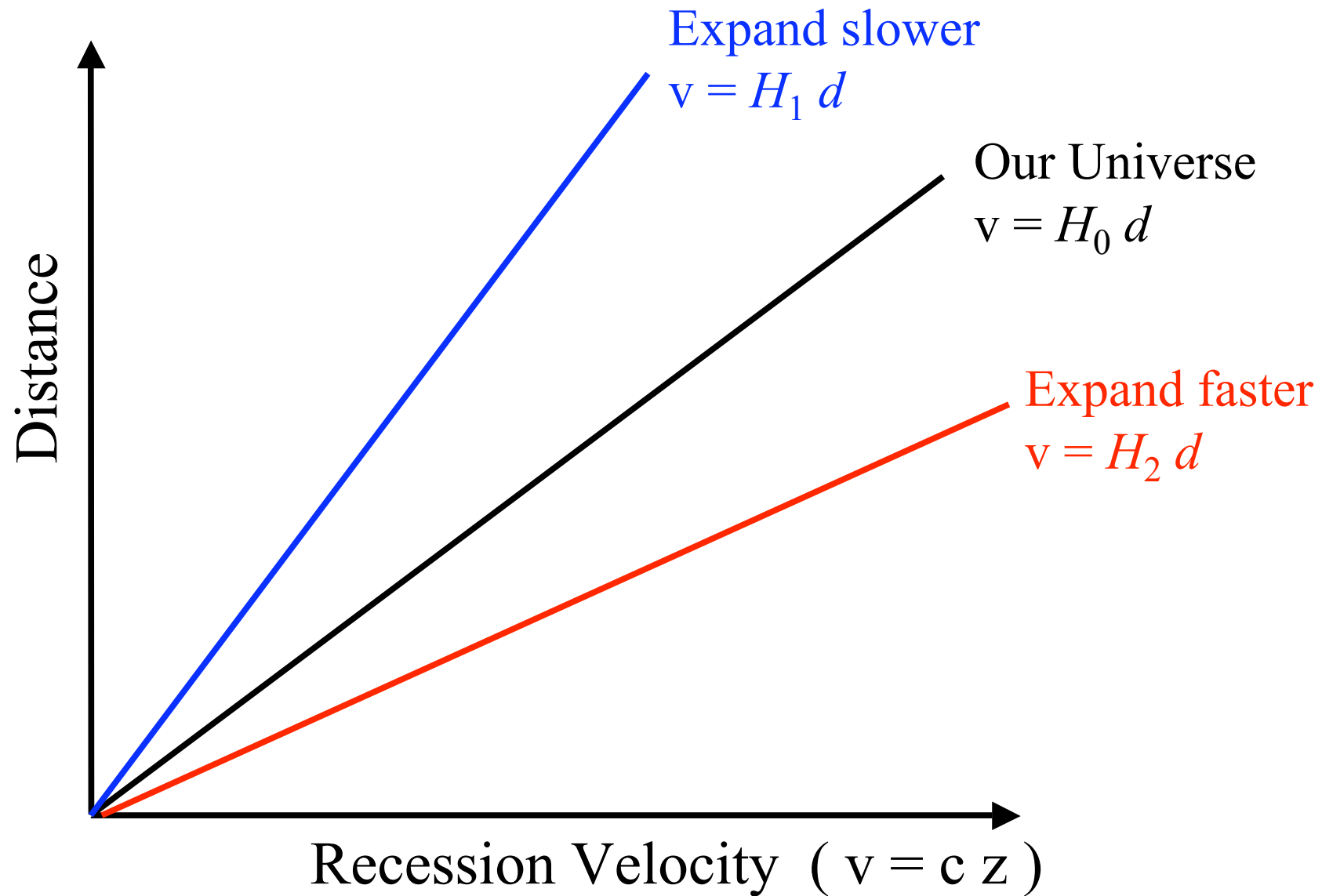
1. Brighter ones decline more slowly.
2. Time runs slower by factor  $(1+z)$ .

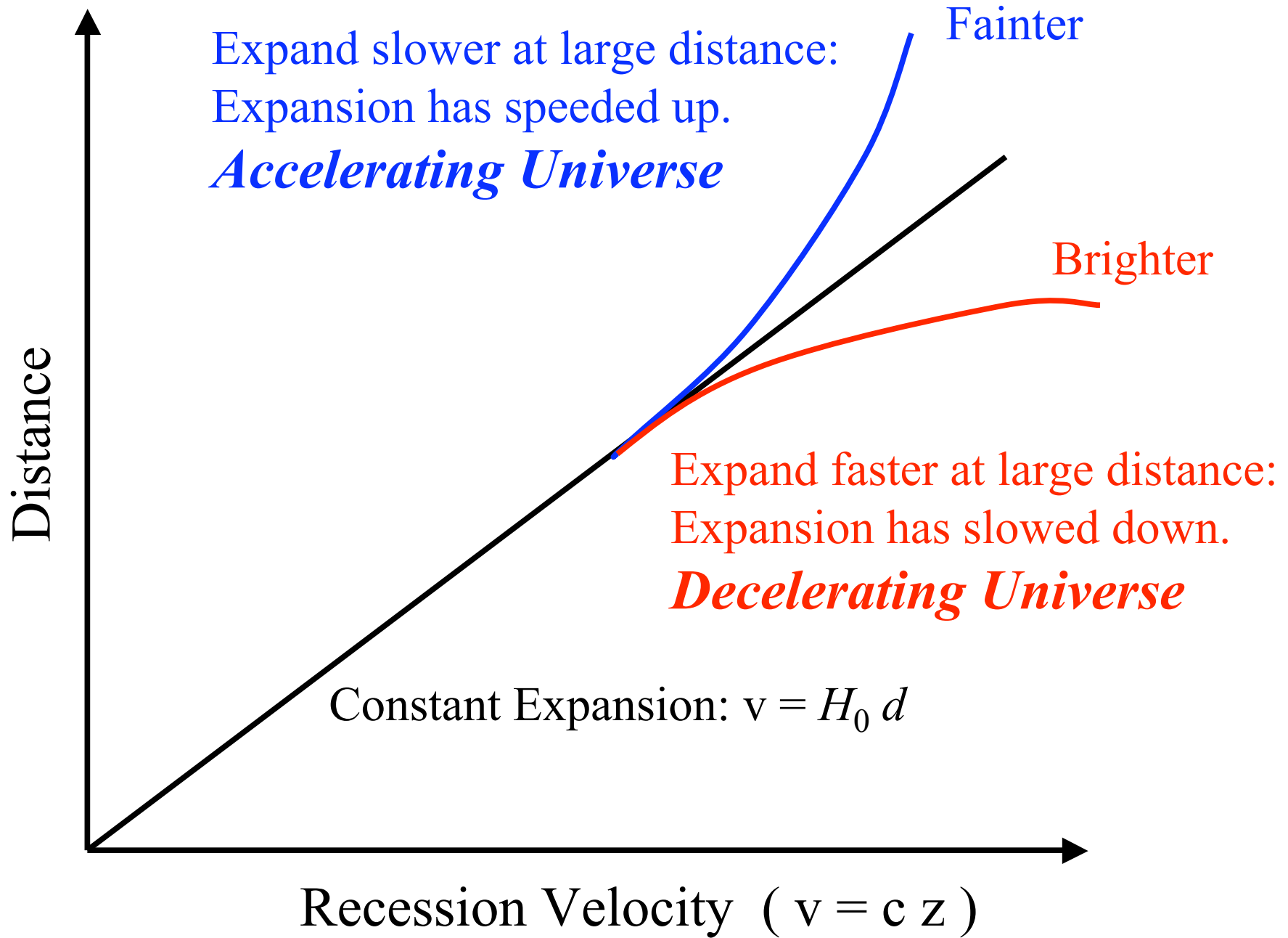
**AFTER correcting:**  
Constant peak brightness  
 $M_B = -19.7$

**Observed peak magnitude:**  
 $m = M + 5 \log (d/\text{Mpc}) + 25$   
gives the distance!

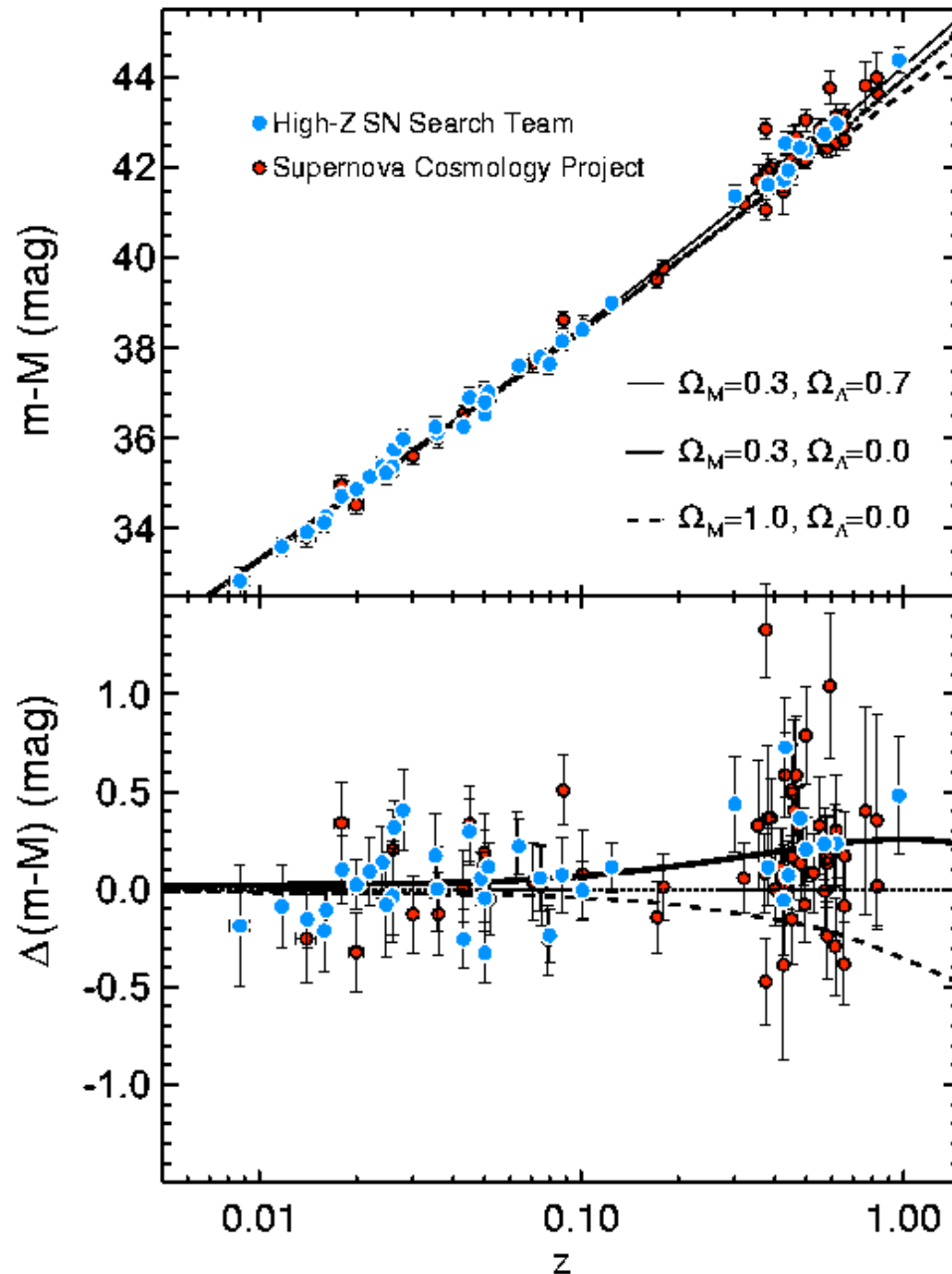


# Varying Hubble Laws





# 1998: Accelerating Universe !?



Distant SNe ( $z \sim 0.8$ ) are **25% fainter** than the Hubble Law.  
 $\Rightarrow$  **ACCELERATING!**

Need cosmological constant  $\Lambda$   
or DARK ENERGY.

Proposed satellite: *SNAP*  
*SuperNova Acceleration Probe*  
to find even more distant SNe ...

# Acceleration by DARK ENERGY

First, gravity from high matter density decelerates the expansion.

Expansion reduces matter density, deceleration slows.

Then, **DARK ENERGY** accelerates.

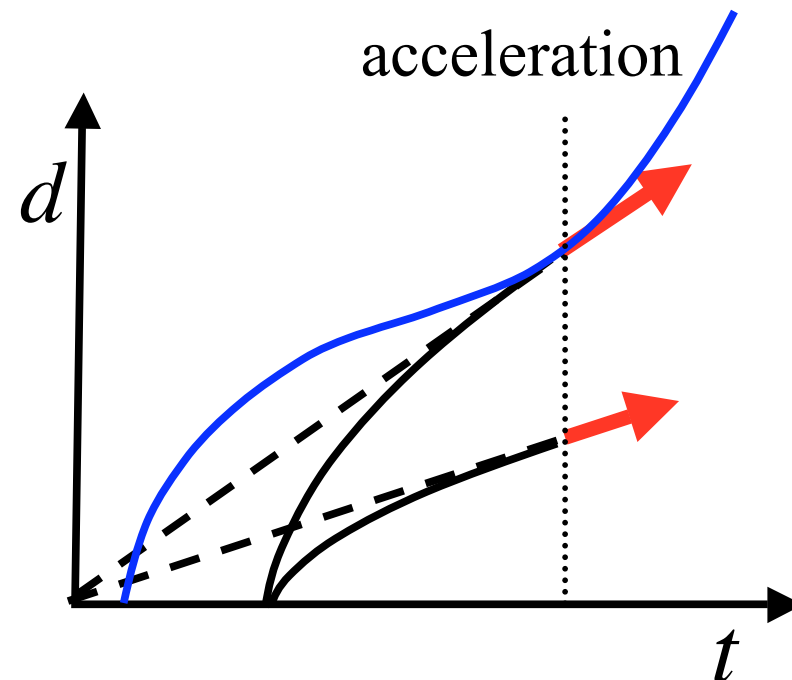
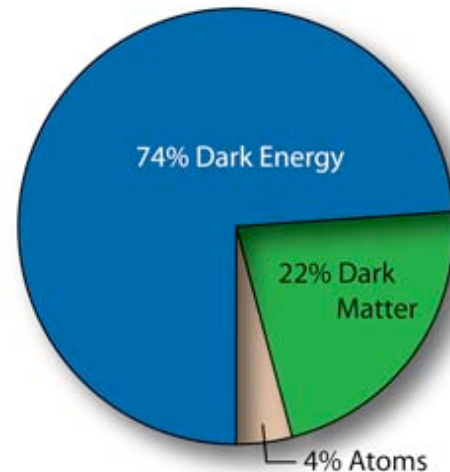
Slight Problem:

Quantum vacuum predicts  
Dark Energy density

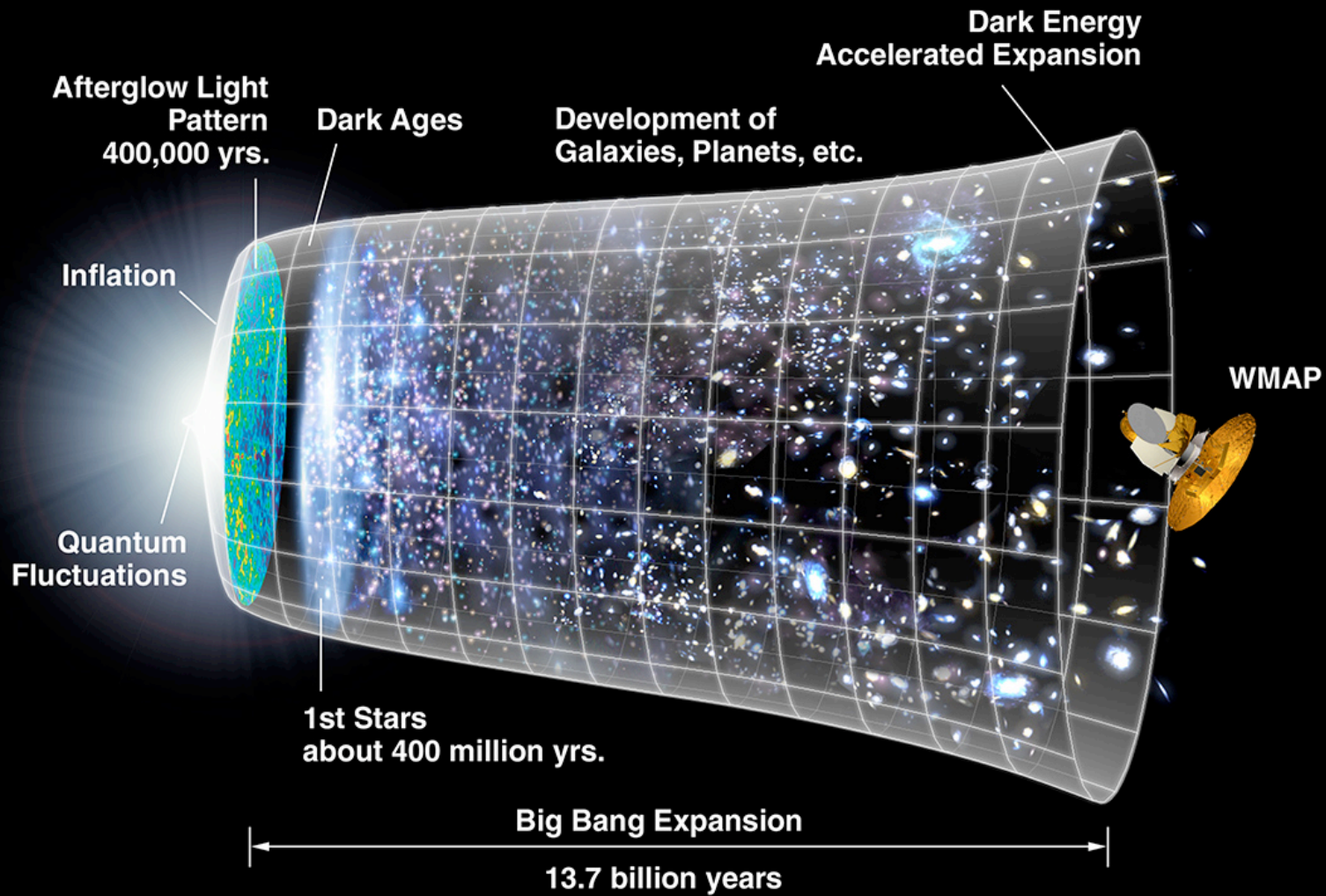
$$\rho_{\Lambda} = 10^{120} \rho_{\text{CRIT}}$$

Observed:

$$\rho_{\Lambda} = 0.7 \rho_{\text{CRIT}}$$

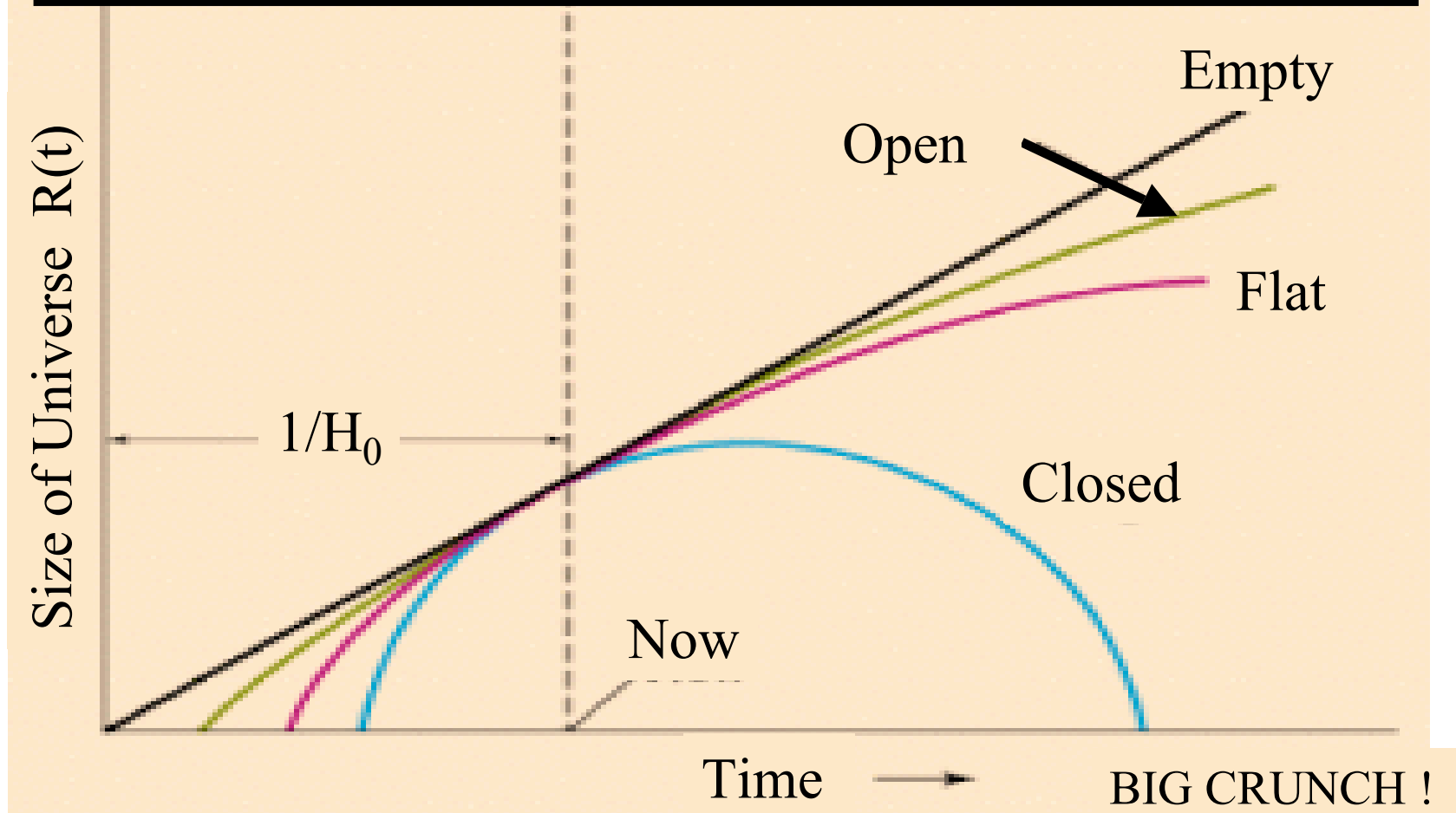






# Re-collapse or Eternal Expansion ?

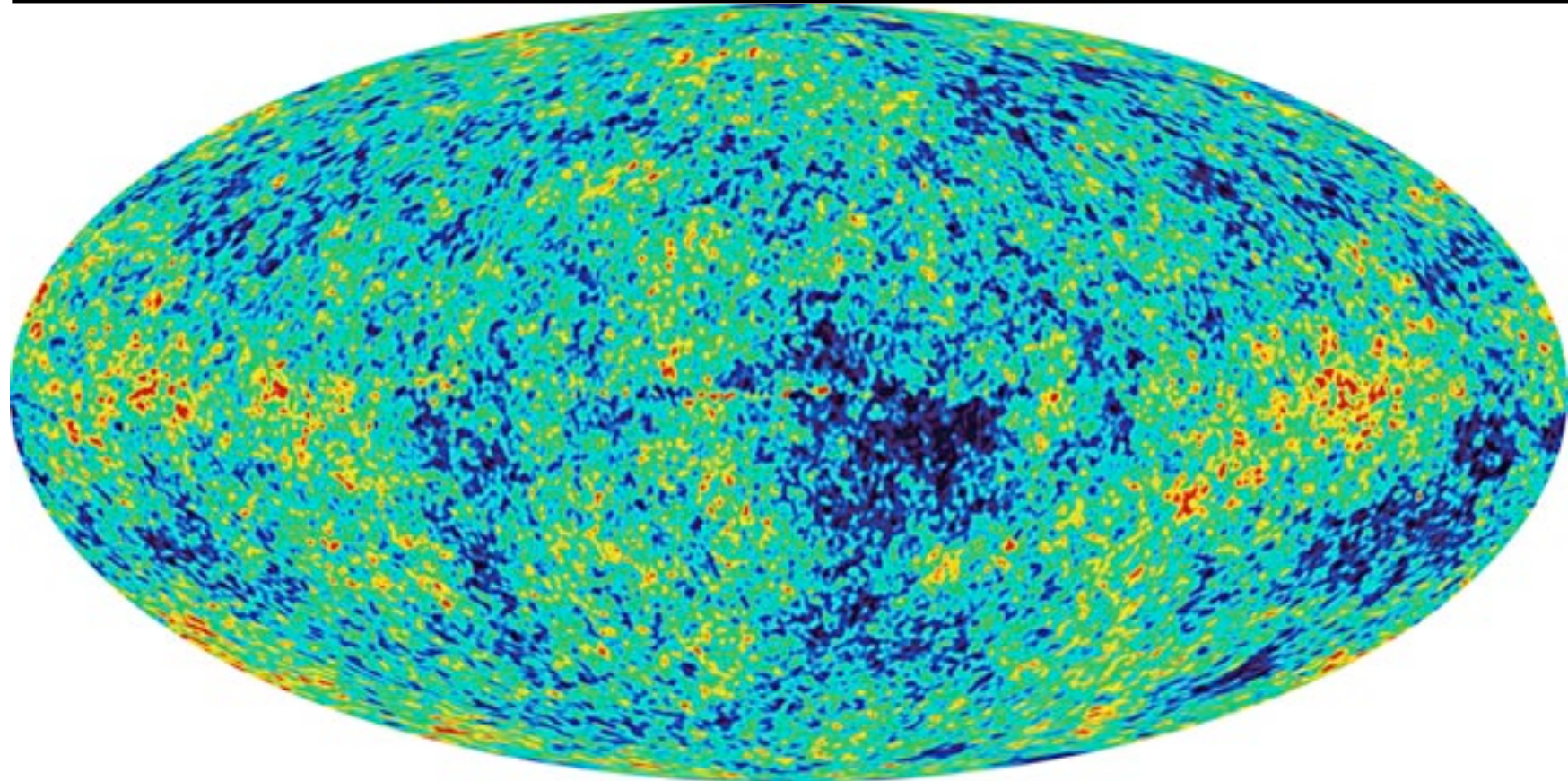
Inflation  $\Rightarrow$  expect **FLAT GEOMETRY**  
**CRITICAL DENSITY**



2004: WMAP all-sky CMB temperature map.

Tiny ripples (at  $z=1100$ ,  $T=3000\text{K}$ ,  $t=3\times 10^5$  yr )  
are the seeds of galaxy formation!

Angular size  $\Delta\theta = 1^\circ \Rightarrow$  FLAT GEOMETRY

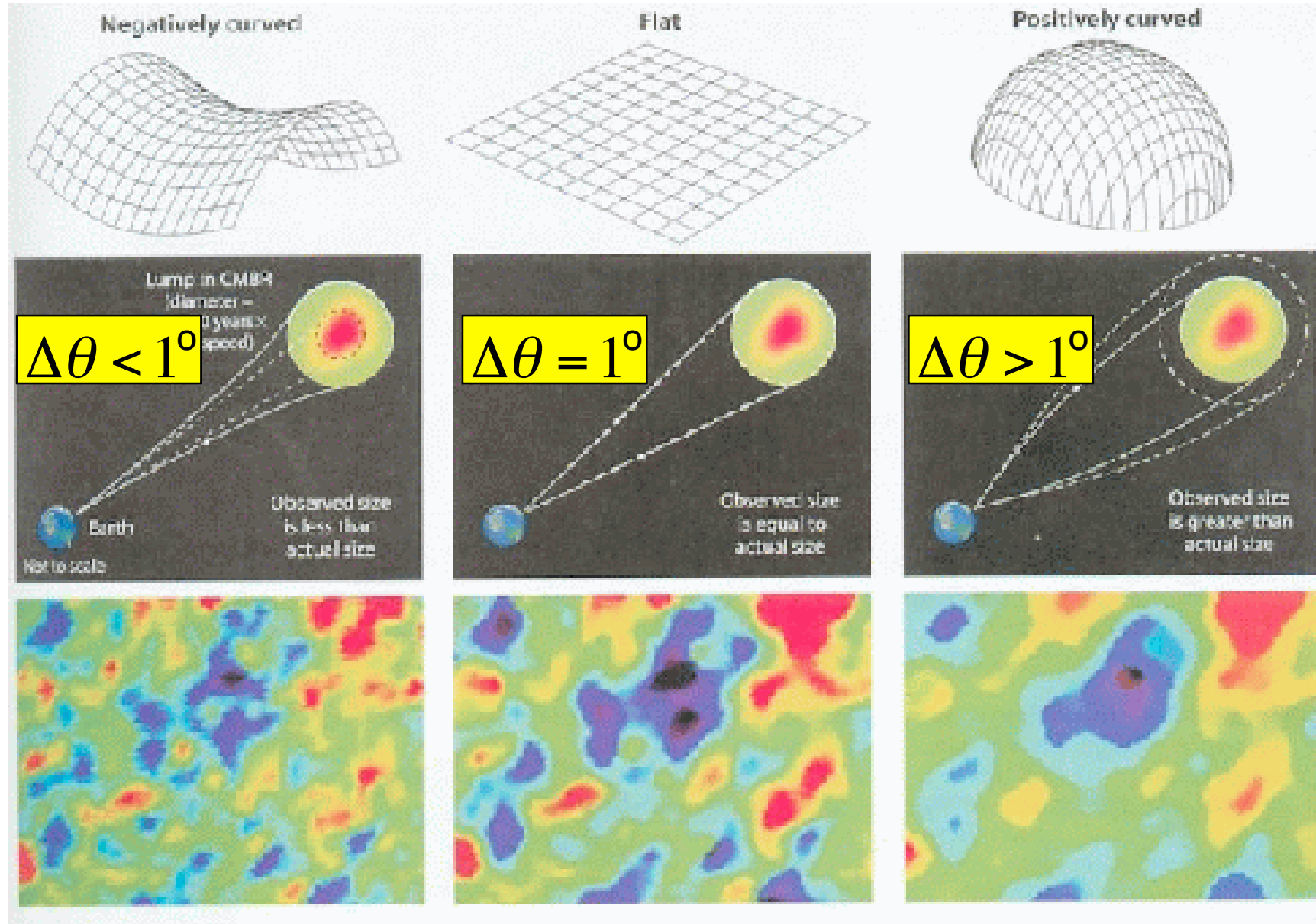


# Curvature of the Universe

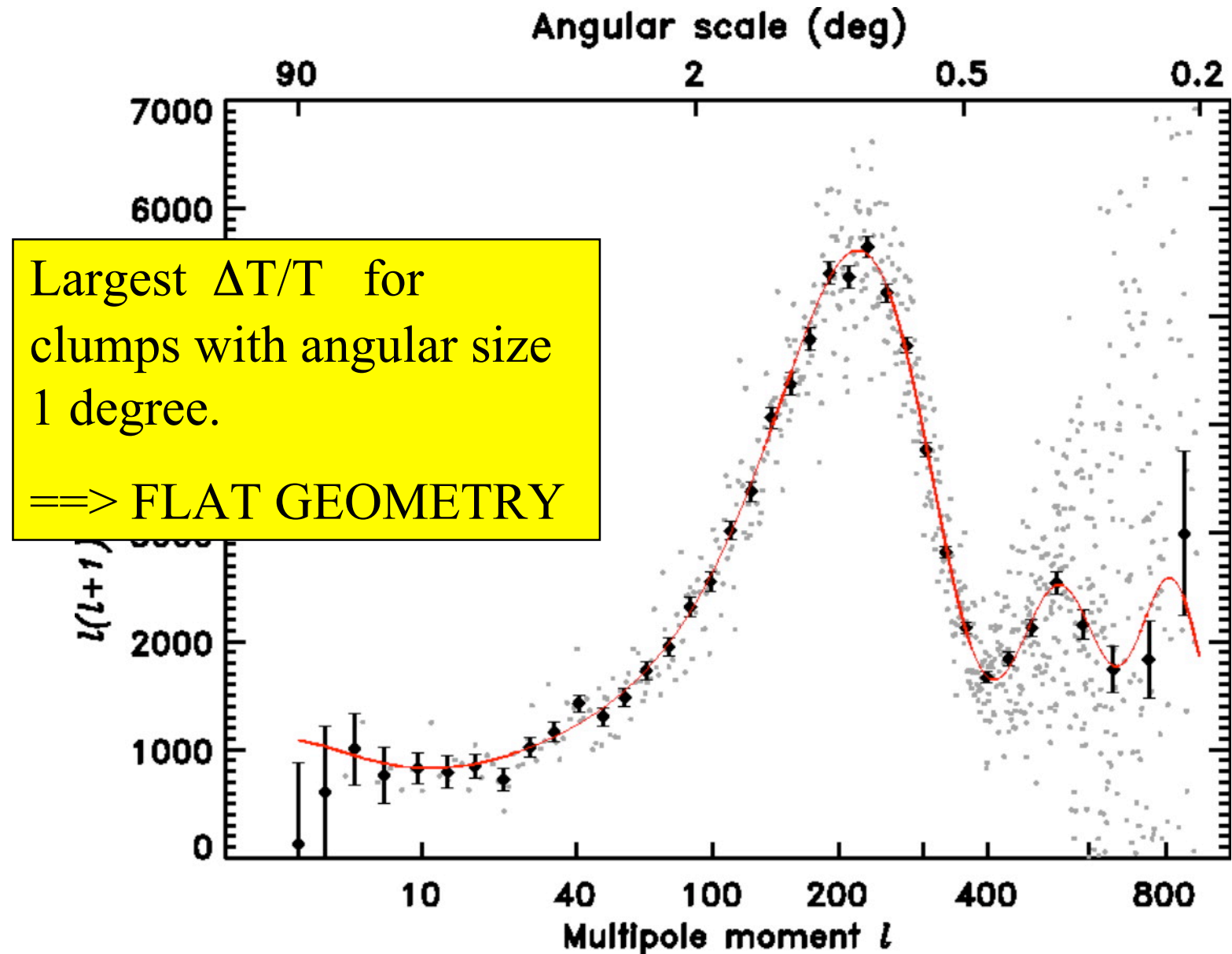
Negative

FLAT

Positive



# 2004: WMAP - Power Spectrum



# Possible Universes

$$H_0 \approx 72 \frac{\text{km/s}}{\text{Mpc}}$$

$$\Omega_M \sim 0.3$$

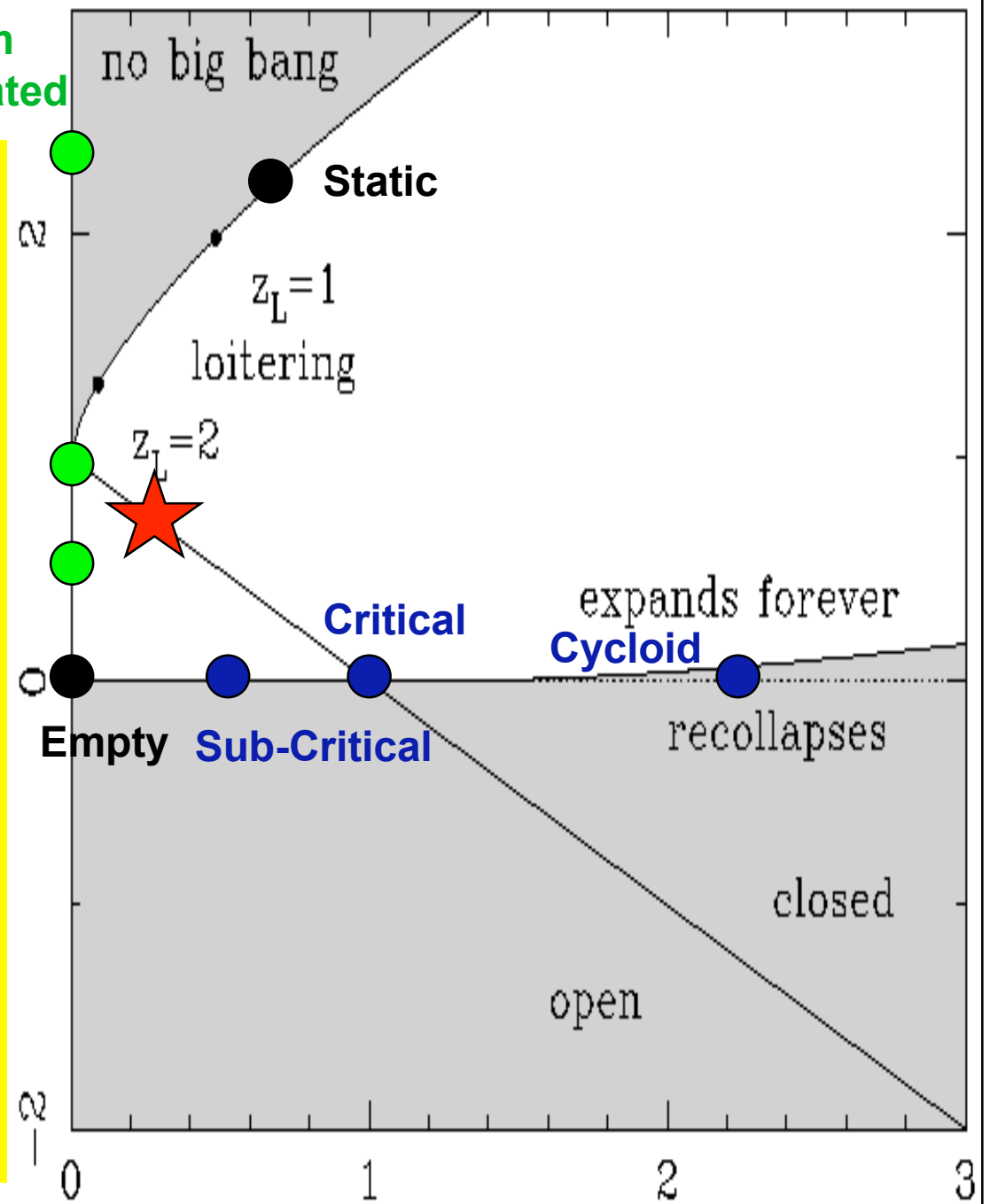
$$\Omega_\Lambda \sim 0.7$$

$$\Omega_R \sim 8 \times 10^{-5}$$

$$\Omega = 1.0$$

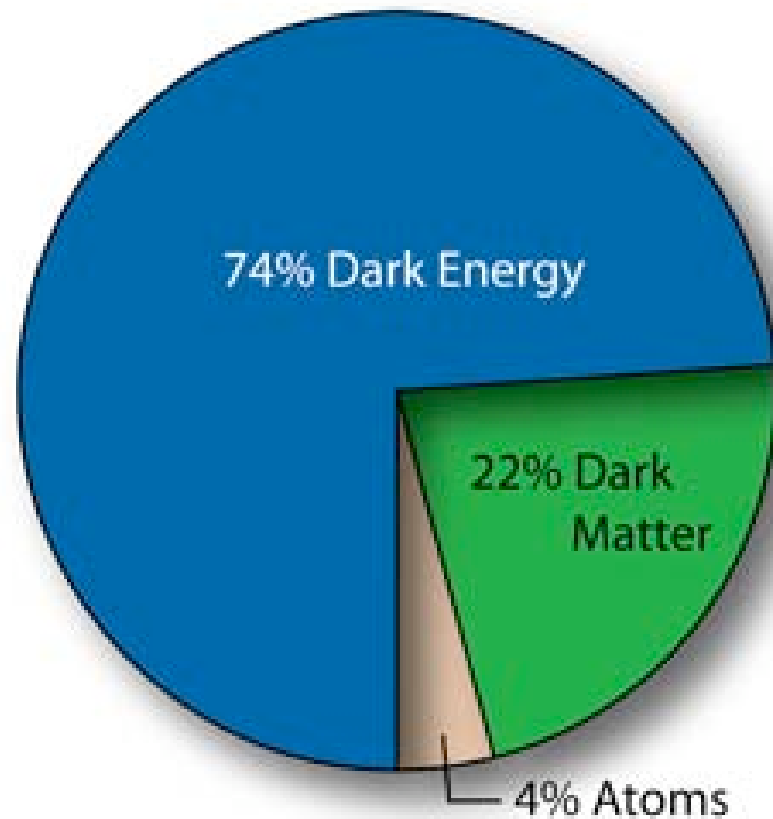
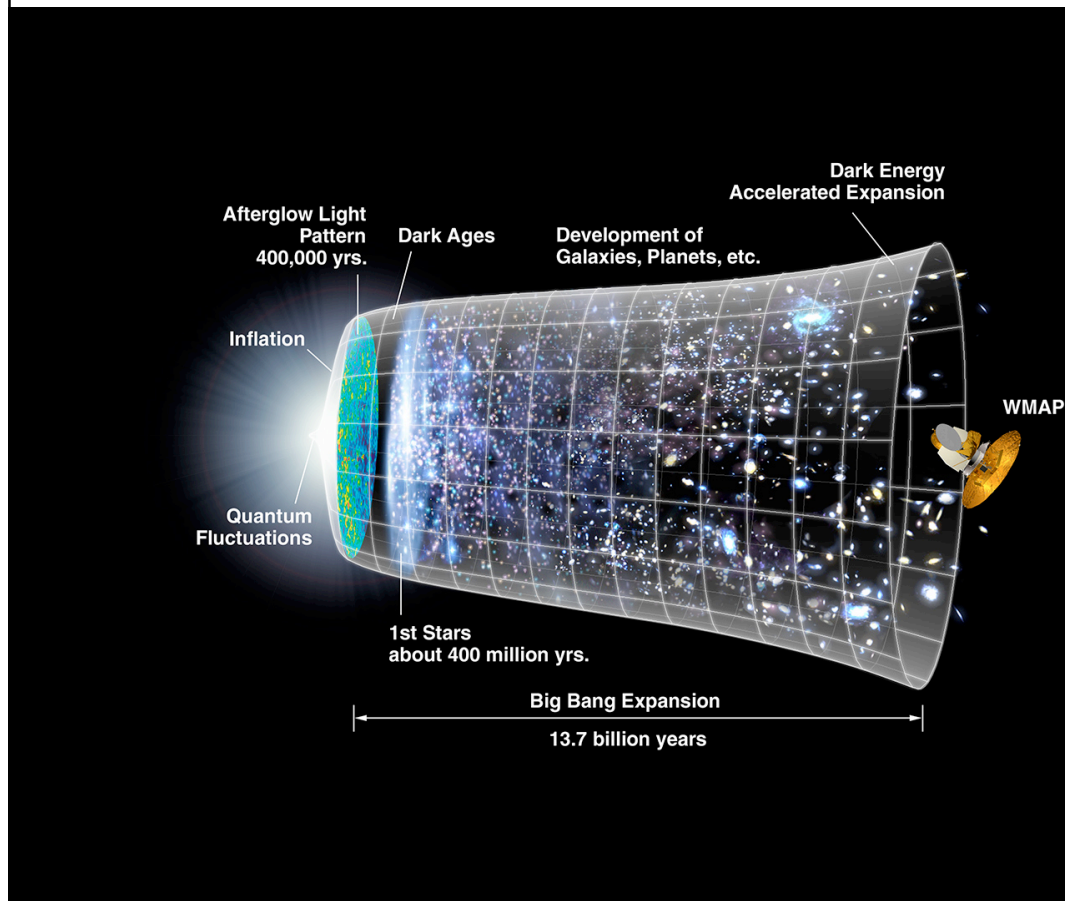
Vacuum  
Dominated

Dark Energy:  $\Omega_\Lambda = \rho_\Lambda / \rho_{\text{CRIT}}$

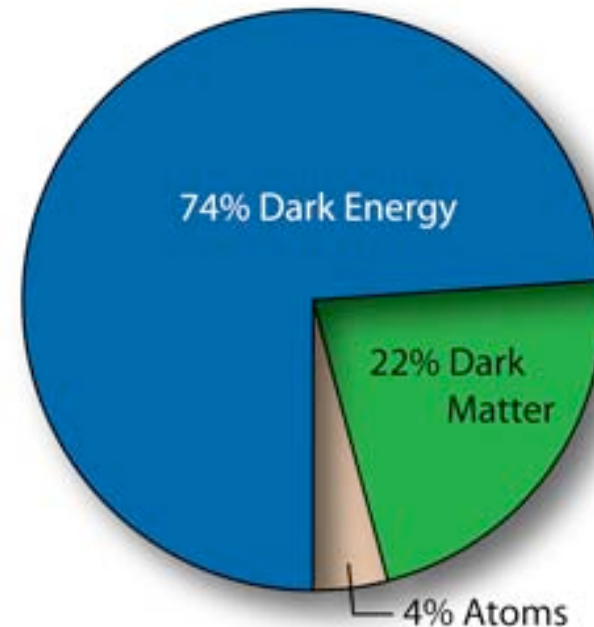


# Our “Crazy” Universe

**~4% Normal Matter**  
**~22% “Dark Matter”**  
**~74% “Dark Energy”**



Or .... Has General  
Relativity Failed ?



Can an ***Alternative Gravity*** Model  
fit all the data without  
**Dark Matter and Dark Energy ?**

**No luck yet, but people are trying.**