

AS1001:Extra-Galactic Astronomy

Lecture 6: Galaxy Orientation, Black Holes & Quasars

Galaxy Inclination

FACE-ON
Inclination $i = 0^\circ$

EDGE-ON
Inclination $i = 90^\circ$

All galaxies are viewed with some inclination:

$b = \text{Minor Axis}$
 $a = \text{Major Axis}$

Calculating the Inclination

- Assuming a thin circular disc:
- Inclination, i , given by: $\cos(i) = \frac{b}{a}$

- $a = b, i = 0^\circ$
- $b = 0, i = 90^\circ$

NB: a is always measurable

Line-of-sight velocity

The **Doppler Shift** measures the component of velocity **along the line of sight**.
Need to correct for inclination.

$i = 0^\circ$
 $V_{OBS} = 0$

$V_{OBS} = V_{ROT} \sin(i)$

$i = 90^\circ$
 $V_{OBS} = V_{ROT}$

Example: Inclination Corrections

A long-slit spectrum aligned with a galaxy's major axis has an [OII] line at 3900A that shifts by 5A from one side to the other side of the galaxy. The major-to-minor axis ratio is 3. What is the rotational velocity of the outermost stars ?

$$\cos i = b/a = 1/3 \quad i = \cos^{-1}(1/3) = 70.5^\circ$$

$$\sin i = \sqrt{1 - \cos^2 i} = \sqrt{1 - (1/9)} = \sqrt{8/9} = 0.94$$

$$V_{OBS} = \frac{\Delta\lambda}{\lambda} c = \frac{2.5\text{A}}{3900\text{A}} \times (3 \times 10^5 \text{ km/s}) = 192 \text{ km/s}$$

Note: (5 A)/2

$$V_{ROT} = \frac{V_{OBS}}{\sin(i)} = 204 \text{ km/s}$$

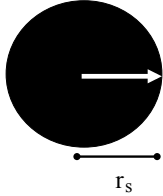
Note: $\lambda=3900$ and not 3727

Black Holes

Gravity = curvature of space-time by matter/energy.
Pack mass into a small enough volume, and the space-time can be so distorted that nothing, not even light, cannot escape.

The Schwarzschild Radius

- Where the escape velocity equals the speed of light.
- Nothing, not even light, can escape from inside the Event Horizon, at the Schwarzschild Radius r_s .
- Escape velocity: set Kinetic Energy = Gravitational Energy



$$\frac{1}{2} m v^2 = \frac{G M m}{r}$$

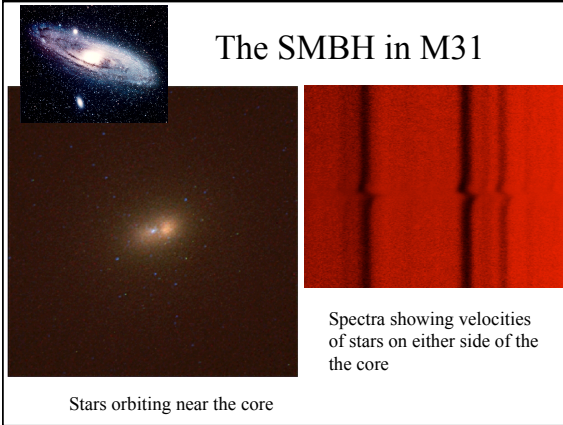
$$v_{\text{esc}} = \left(\frac{2GM}{r} \right)^{1/2}$$

$$r_s = \frac{2GM}{c^2} = 3 \text{ km} \left(\frac{M}{M_{\text{sun}}} \right)$$

Types of Black Hole

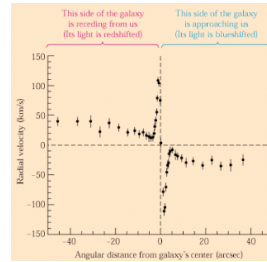
- Stellar-mass
 - Formed when a very massive star goes supernova
$$M_{\text{BH}} \sim 10 M_{\odot}$$
- Super-massive
 - Formed in galaxy cores
$$M_{\text{BH}} \sim 10^{7-9} M_{\odot}$$
- Most large galaxies have a super-massive black-hole (SMBH) in their core.

The SMBH in M31



The SMBH in M31

- Star velocities $v > 110 \text{ km/s}$ for $r < 2.5 \text{ pc}$



Use Virial Theorem to calculate the Mass

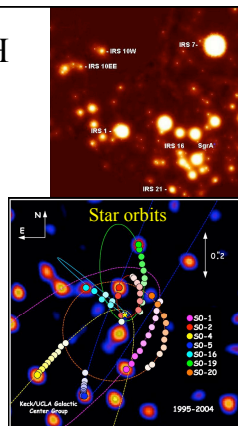
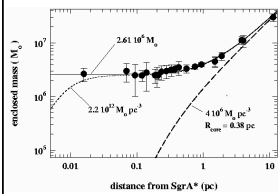
$$M_{\text{CORE}} = \frac{v^2 r}{G}$$

$$= \frac{(110 \times 10^3)^2 \times 2.5 \times (3 \times 10^{16})}{6.67 \times 10^{-11}}$$

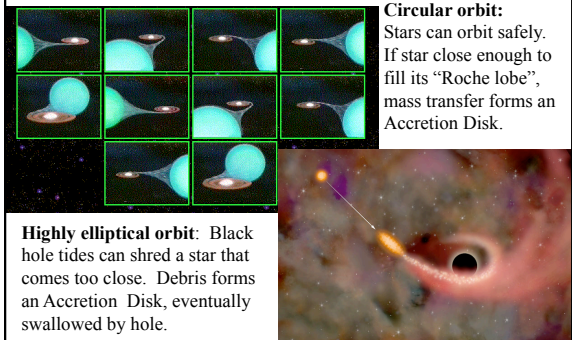
$$= 1.4 \times 10^{37} \text{ kg} = 6.8 \times 10^6 M_{\odot}$$

Milky Way's SMBH

- Infrared images (to see thru dust) show a compact star cluster.
- Star velocities $v > 1000 \text{ km/s}$ inside $r = 0.01 \text{ pc}$!
- $\Rightarrow 3 \times 10^6 M_{\text{sun}}$ Black Hole in the Milky Way's core

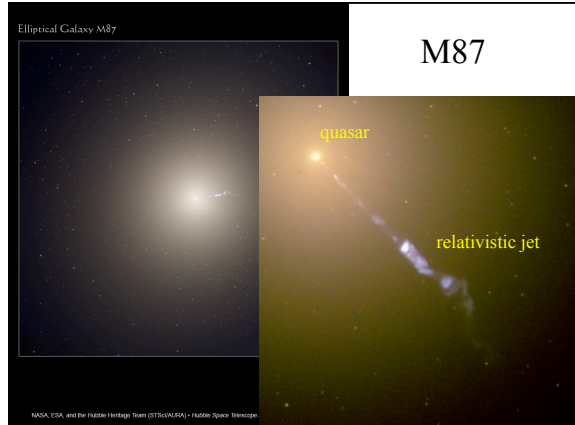


Stars orbiting a Black Hole



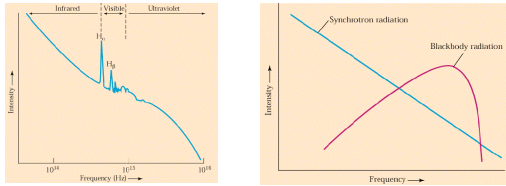
Discovery of Quasars

- Quasars are Super-Massive Black Holes “feeding”.
- Originally known as Quasi-Stellar Objects (QSOs).
- For many years “stars” with unknown spectral features were found but their nature unknown !
- 1963: Martin Schmidt recognised that QSOs have known emission lines with large redshifts (hence QSO luminosities >> galaxies)
- Hence QSOs are extra-galactic objects but:
 - Appear star-like (i.e., not extended but point-like)
 - Outshine galaxies (by up to 10^5 times)
 - Very broad emission lines ($\Delta v \sim 10^4$ km/s)
 - X-ray and radio emission (from relativistic jets)



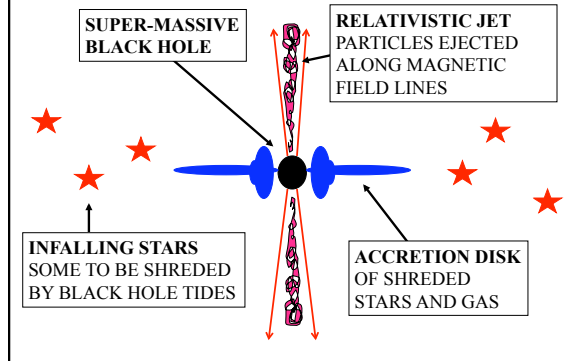
Quasar Spectra

- QSO spectra show both thermal (Blackbody) and non-thermal (Synchrotron) emission.

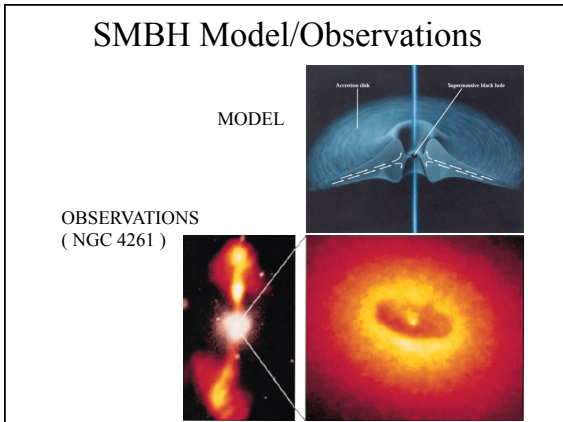


- Blackbody from multi-temperature Accretion Disk.
- Synchrotron from Relativistic Jets: relativistic charged particles spiraling around magnetic field lines.

Simplified Quasar Model



SMBH Model/Observations



Quasars: Powered by Accretion

- Gravitational energy is released as mass accretes.
- Friction in the accretion disk moves angular momentum outward as the gas spirals in. Friction also heats the gas.

• Accretion Disk Temperature Profile:

$$T(r) \sim \left(\frac{3GM\dot{M}}{8\pi\sigma r^3} \right)^{1/4} = 10^6 K \left(\frac{\dot{M}}{M_{\text{SUN}}/\text{yr}} \right)^{1/4} \left(\frac{M}{10^8 M_{\text{SUN}}} \right)^{-1/2} \left(\frac{r}{r_s} \right)^{-3/4}$$

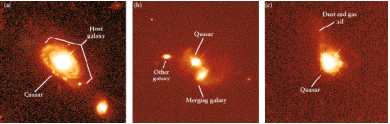
- Accretion Luminosity:

$$L \sim \frac{GM\dot{M}}{r_s} = \eta \dot{M} c^2 \sim 10^{11} L_{\text{SUN}} \left(\frac{\dot{M}}{M_{\text{SUN}}/\text{yr}} \right)$$

- η = Efficiency of converting rest mass energy into light:
 - Up to 15% for accretion onto a black hole
 - Much smaller for nuclear fusion

Types of Active Galactic Nuclei

- 1993: HST reveals "Quasar fuzz" = host galaxy.

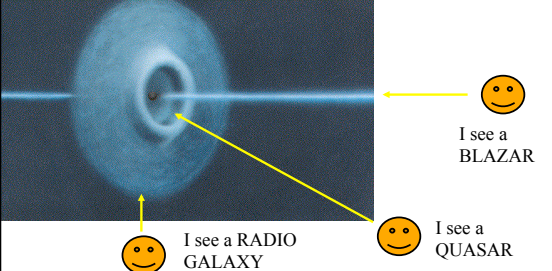


- Active Galactic Nuclei (AGN) in the cores of galaxies.
 - Quasar = Bright AGN outshining the host galaxy
 - Seyfert = Fainter AGN luminosity equals host galaxy
 - Radio Galaxy = AGN with radio lobes
 - Blazar = AGN with no lines and rapid variability

WHY SO MANY DIFFERENT TYPES ?

AGN Unification

Different AGN types are now understood to be due to different viewing angles:



The Quasar Era

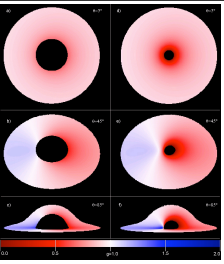
- Redshift surveys find highest density of quasars per unit volume around redshift $z \sim 2-3$
- Large redshifts \Rightarrow large distances \Rightarrow large "lookback times" i.e., we see quasars as they were in the past, when the Universe was young.
- Nearest quasar: 3c273 at 250 Mpc vs 5 Mpc typical galaxy-galaxy distance.
- Thus, very low density of quasars today.
- Quasars were once common, but then died out.

Quasars and Galaxy Formation

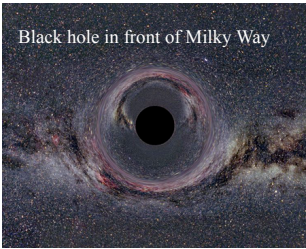
- All large nearby galaxies harbour a SMBH. When swallowing stars/gas, this becomes an AGN.
- During a mad feeding frenzy (e.g. triggered by merger with another galaxy) the SMBH may eat 1000 Msun/year. The galaxy temporarily becomes a Quasar, with an AGN 1000 times brighter than the starlight from the galaxy.
- SMBH, AGN activity, and Quasars are important for galaxy formation (e.g. SMBH mass is always a few % of the stellar bulge mass) but full story still being worked out.

Black Hole Lensing Effects

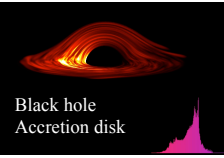
- *Light rays bend toward the mass.
- *Emerging photons are redshifted.
- *Light is beamed and boosted in the direction of relativistic motion.



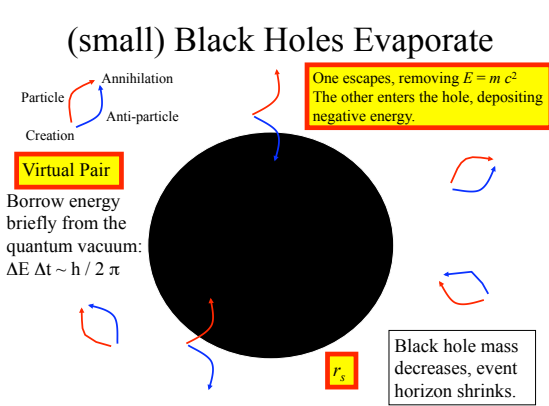
Black hole in front of Milky Way



Black hole Accretion disk



(small) Black Holes Evaporate



Annihilation
Particle
Anti-particle
Creation

Virtual Pair

Borrow energy briefly from the quantum vacuum:
 $\Delta E \Delta t \sim h / 2 \pi$

One escapes, removing $E = m c^2$
The other enters the hole, depositing negative energy.

Black hole mass decreases, event horizon shrinks.

Black Holes Evaporate

- **Hawking Radiation:** involves gravity (G), relativity (c), thermodynamics (k) and quantum mechanics (h).
- Black holes emit Blackbody radiation with a temperature

$$kT = \frac{hGM_{bh}}{4\pi^2 cr_s^2} = \frac{hc^3}{16\pi^2 GM_{bh}}$$

$$L_{bh} = 4\pi r_s^2 \sigma T^4$$

- Luminosity:
- Energy available: $E_{bh} = M_{bh}c^2$
- Evaporation time: $t_{bh} = E_{bh} / L_{bh}$
- $M_{bh} = 10^{15}$ kg (Everest) : $t_{bh} = 15$ billion years
- $M_{bh} = 5 M_{\text{sun}}$: $t_{bh} = 10^{62}$ years